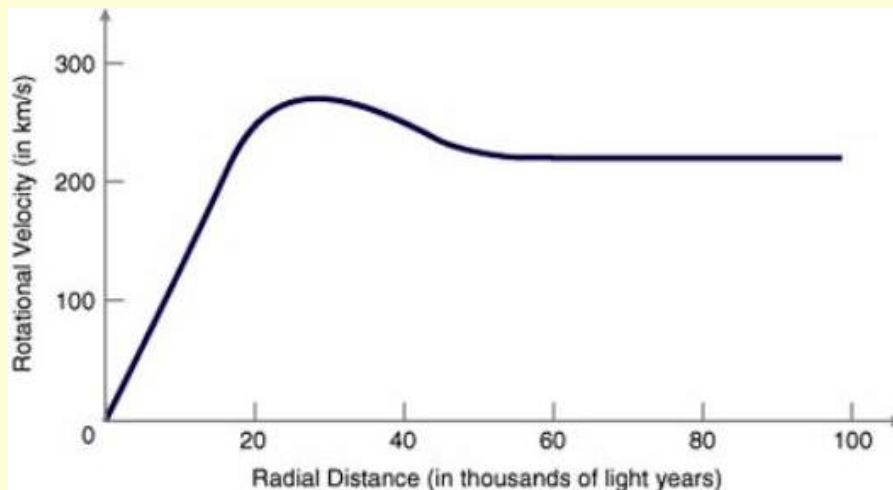


# Search for WIMPs

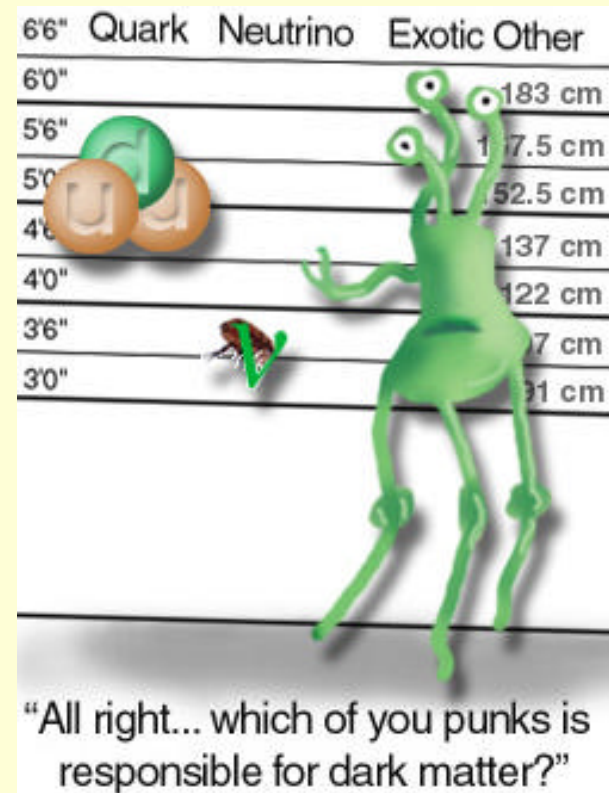
Beatriz E. Burrola Gabilondo

April 19, 2007

# The dark matter problem




<http://www.geocities.com/newastronomy/Chapter13.htm>




[http://physics33.creighton.edu/faculty/duda/dark\\_matter.htm](http://physics33.creighton.edu/faculty/duda/dark_matter.htm)



# What are WIMPs?

- Weakly Interacting Massive Particles
  - Massive – because we have not seen them in our particle accelerators yet
  - Weakly interacting – in agreement with galaxy formation
  - Most popular candidate:
    - Neutralinos
- 




# How do we plan on detecting them?

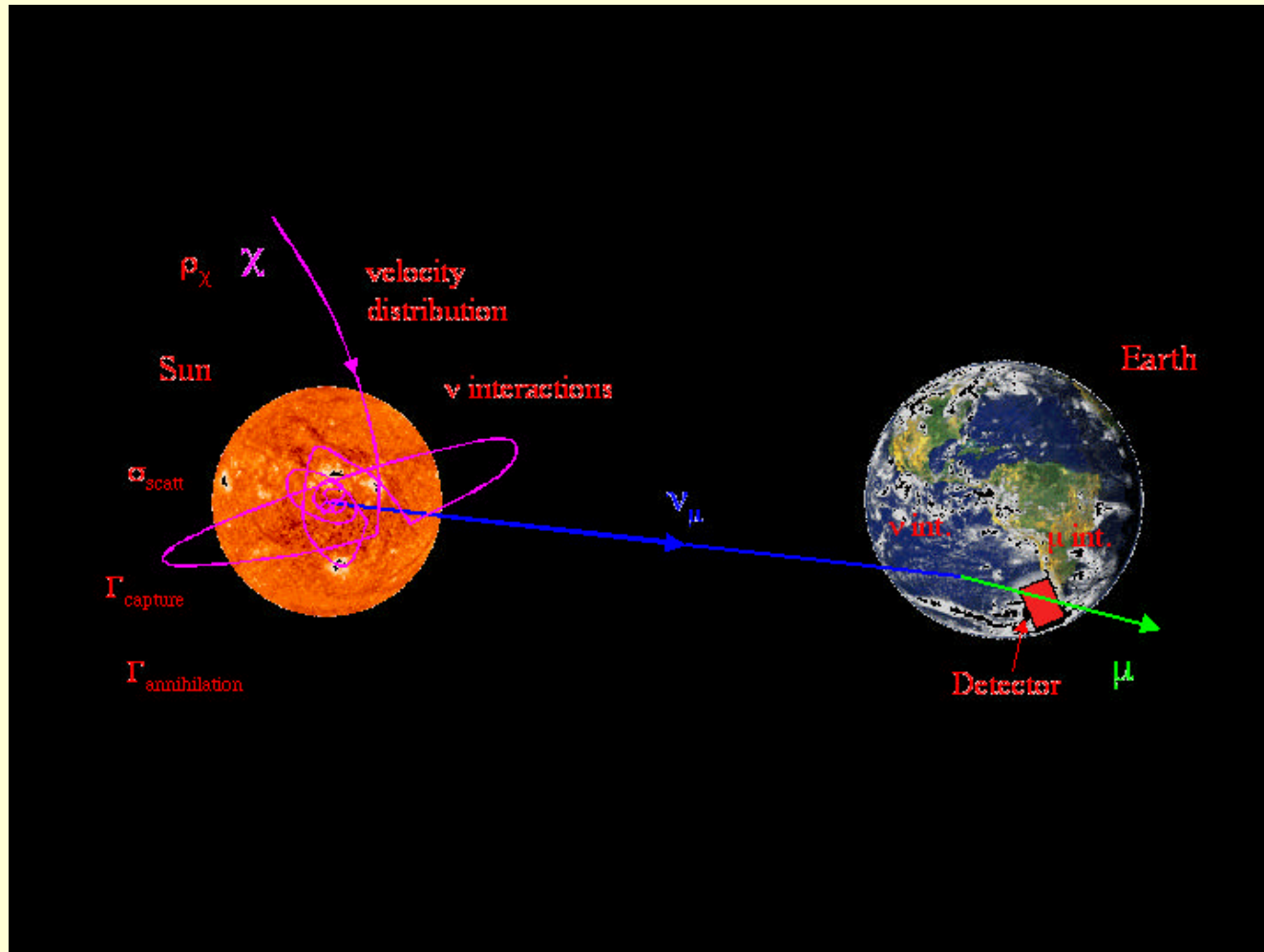
## Indirect methods

- Detect products (neutrinos, gammas) of WIMP decay that happens in the sun

## Direct methods

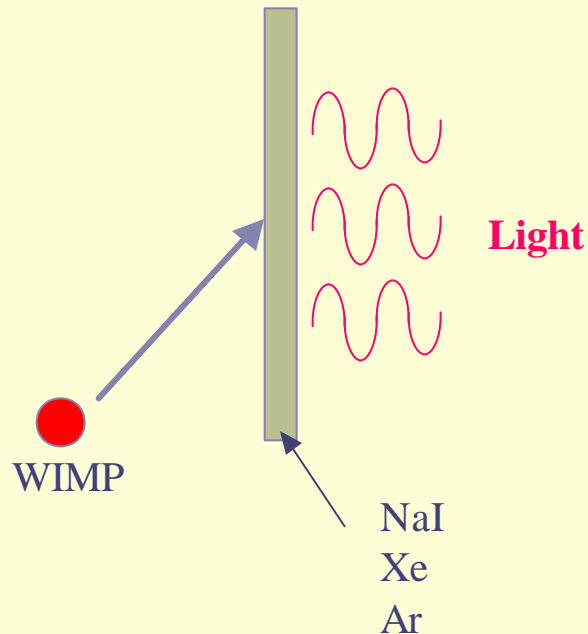
- Detect signature WIMP interaction with matter, or
  - Detect signature annual modulation due to WIMP halo presence
- 

# Solar capture

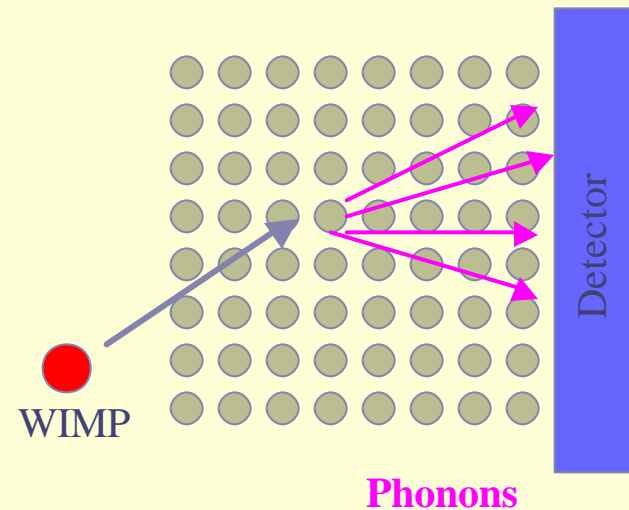


# Direct detection methods

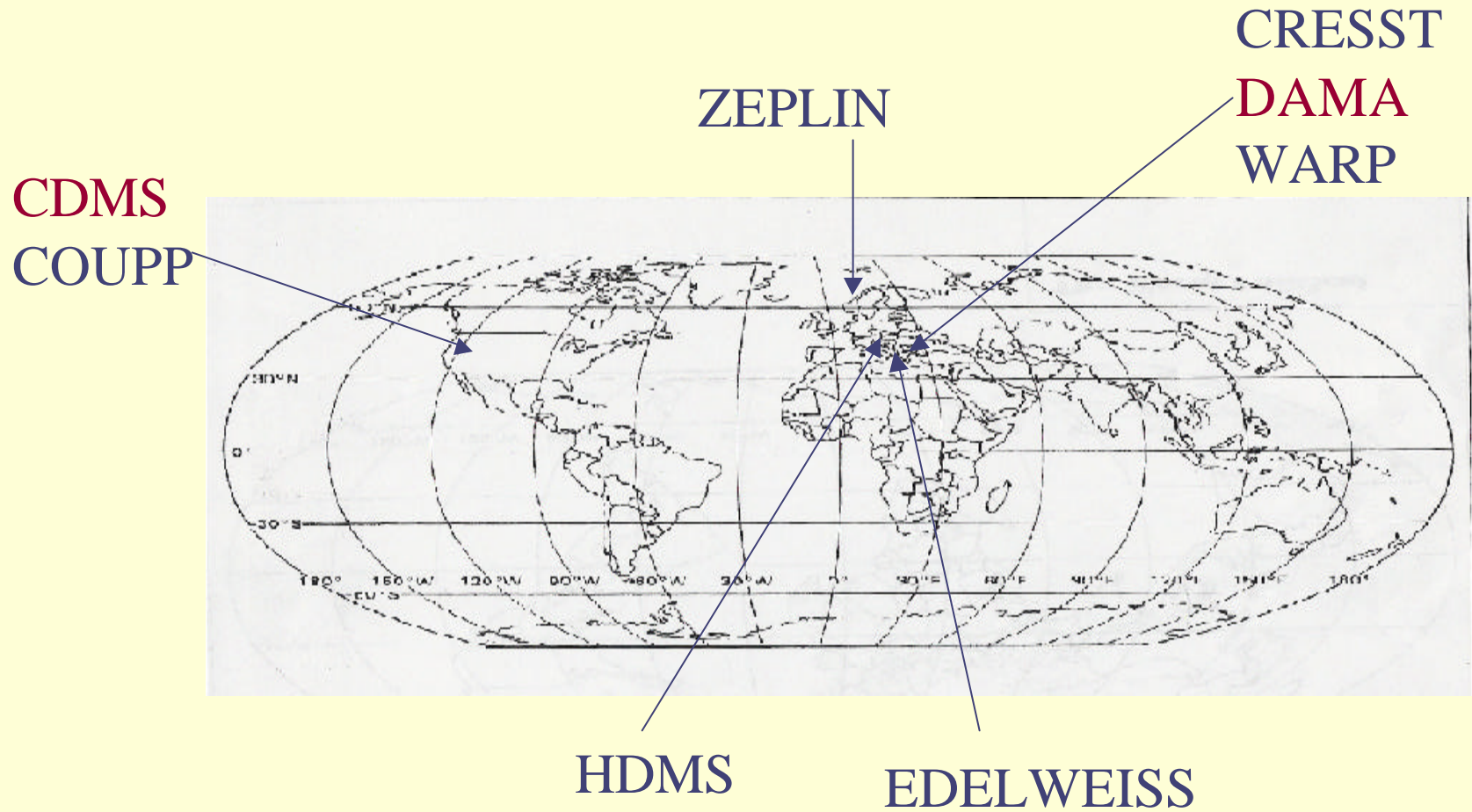
## ● Scintillating materials



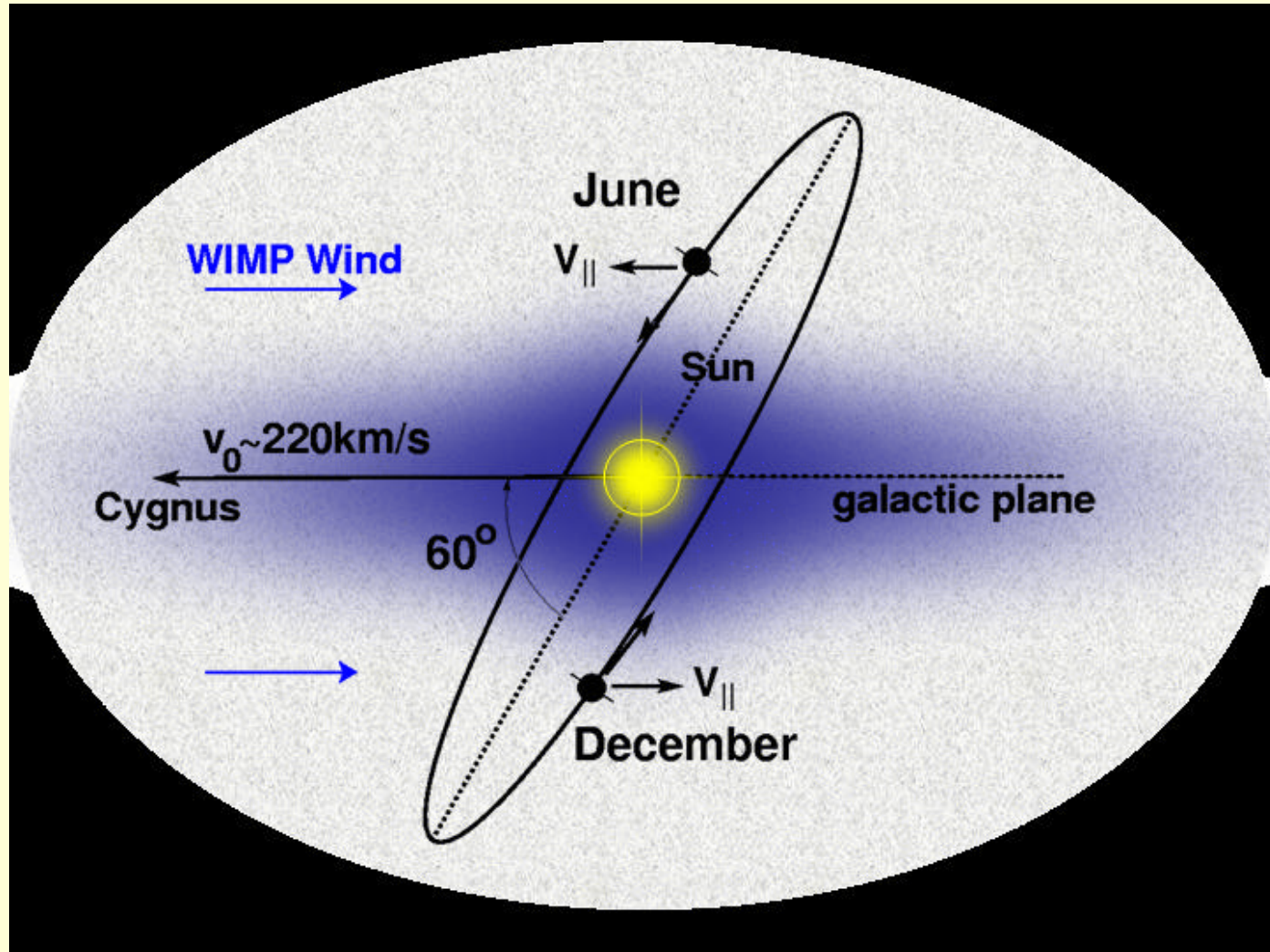
## ● Phonons in Ge or Si



# Experiments around the world



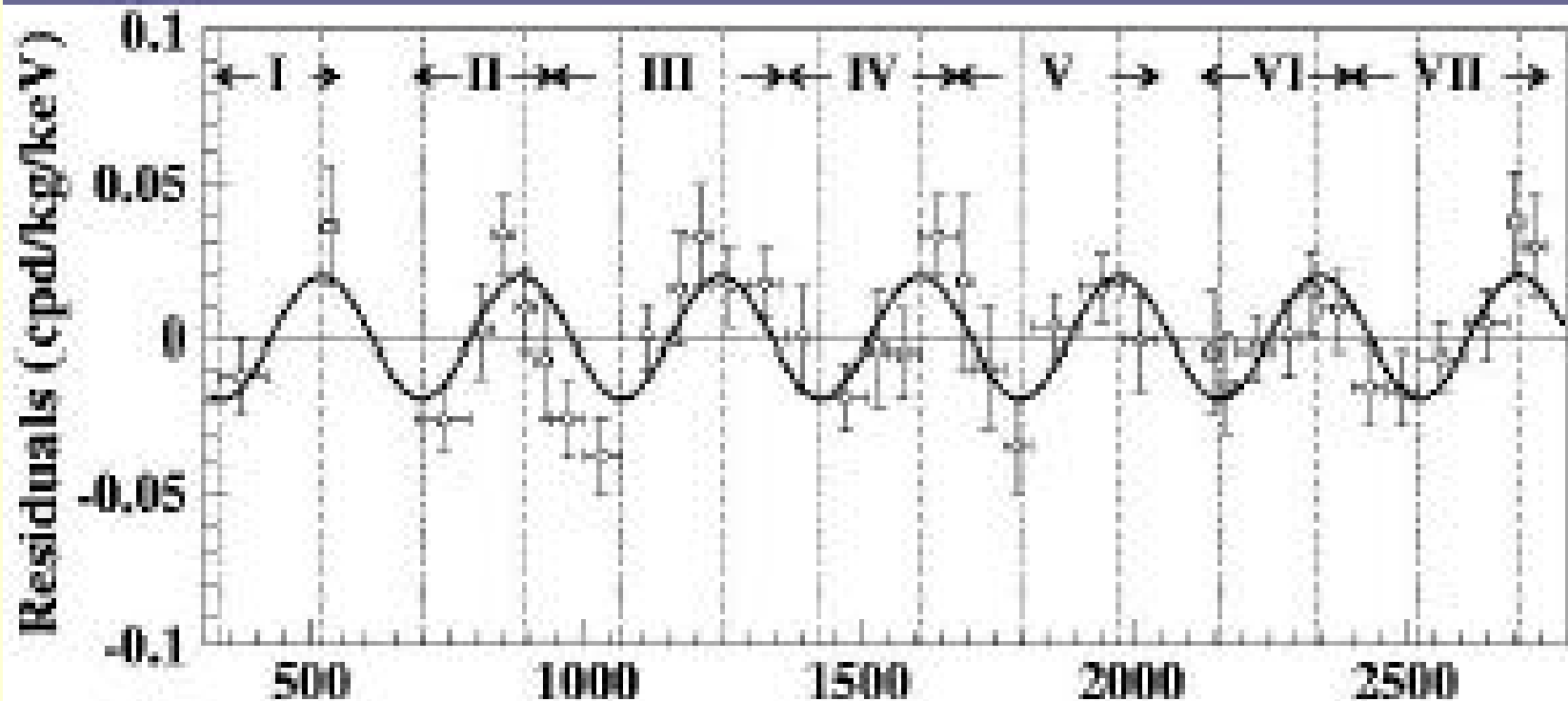
# Annual modulation



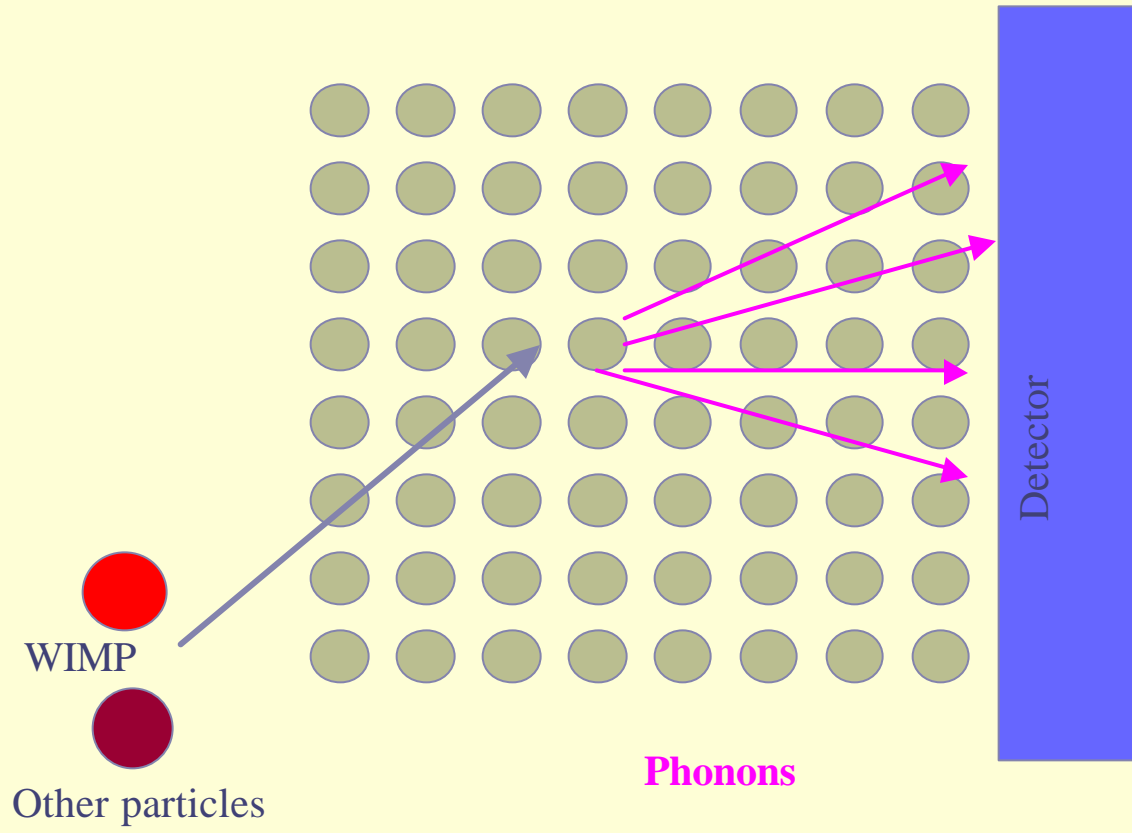


# DAMA results

Experimental residual rate of the single hit events in 2-6 keV over 7 annual cycles

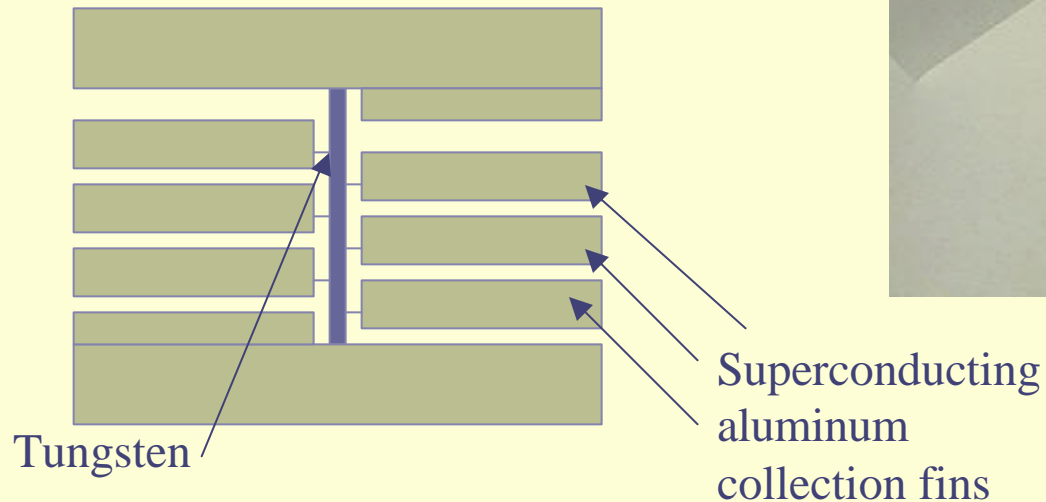
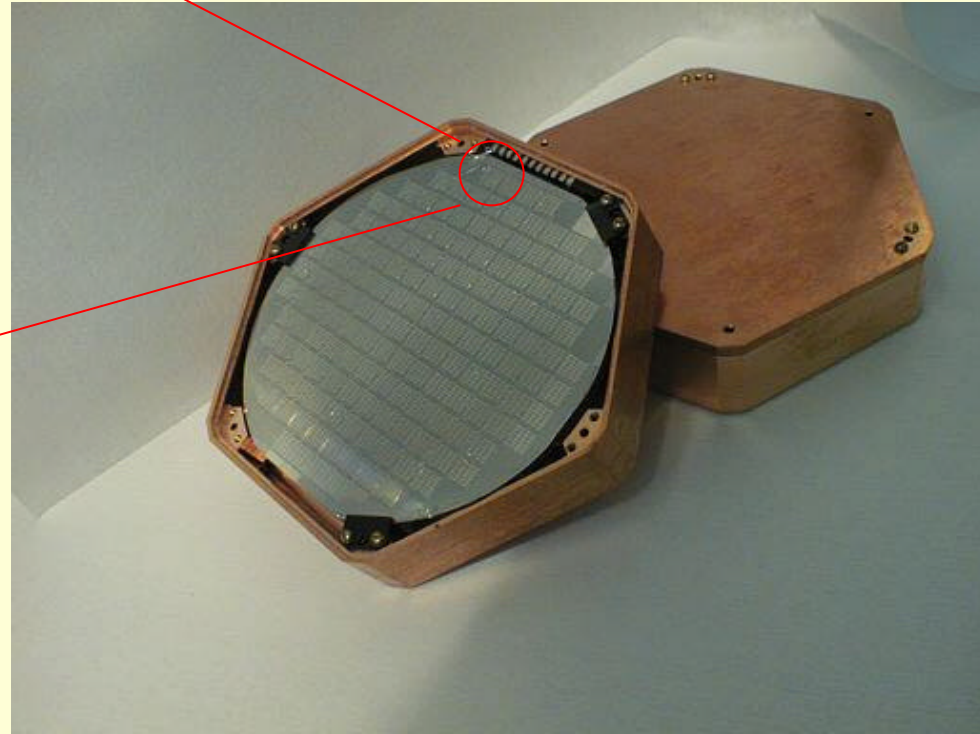
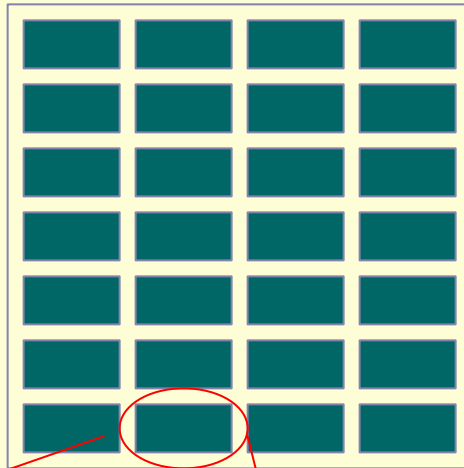


# CDMSII



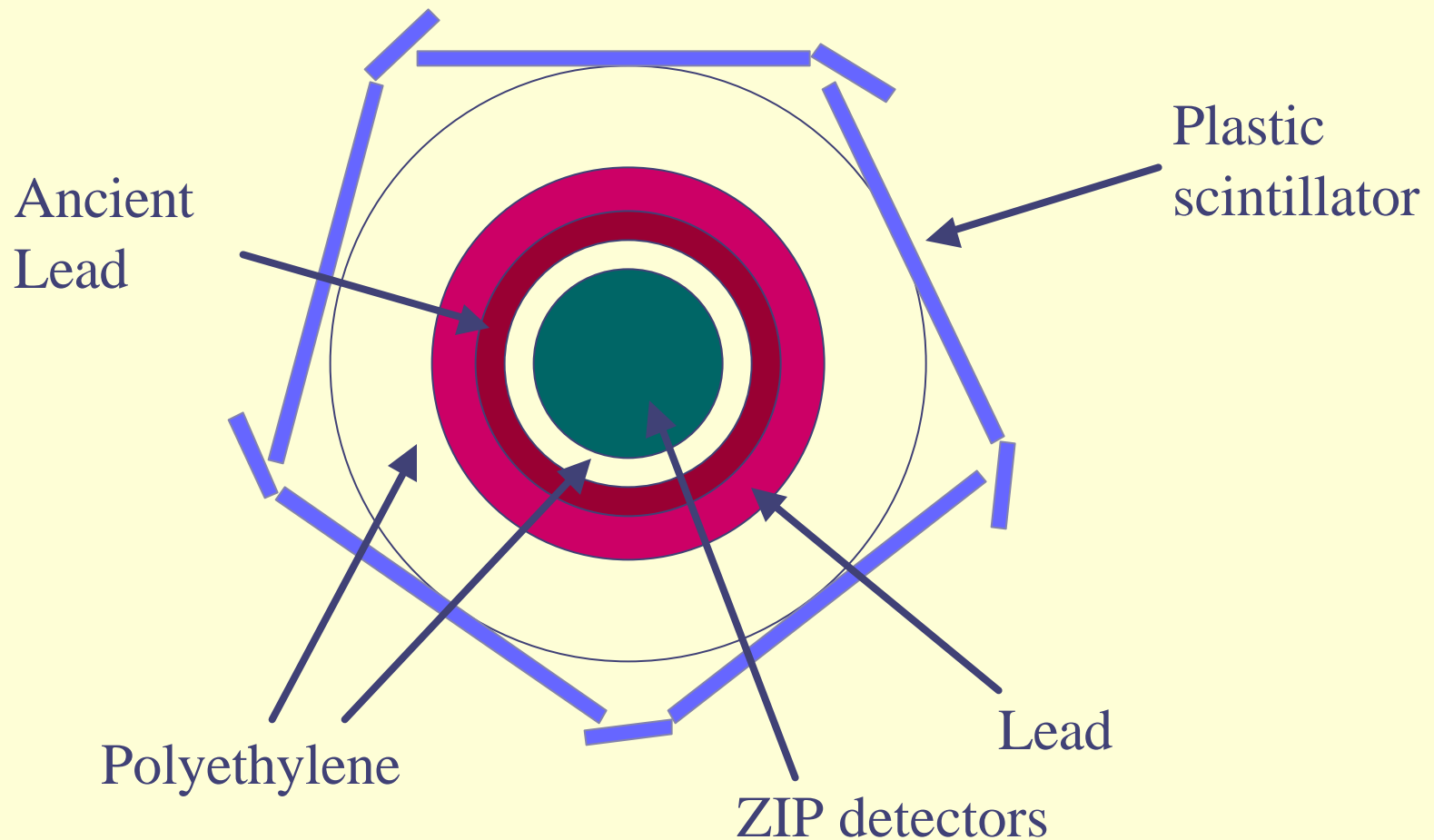
# The CDMSII detectors

148 arrays of  
28 QETs each



Ge or Si ZIP detector

# The CDMSII veto system



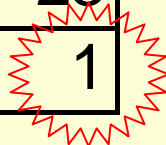


# Results (2004)

All events	968,680
Not random trigger	940,619
Analysis thresholds	79,460
Singles	20,907
Data quality	19,027
Pile up	17,793
Muon veto	17,339
Ionization threshold	14,835
Fiducial volume	7,615
Nuclear-recoil band	23
Phonon timing	1

Only one event in one of the 6 detectors could be a WIMP signal.

Unfortunately, this event is also consistent with expected misidentification due to electron recoils near the surface.



# Where do we stand now?

