

Lecture 29, Nov. 6 (Fri.): part II

Outline for Coulomb potential
(one-electron atom)

- solving radial part of wavefunction, thus energy values
- properties of energy levels
- onto spin-orbit interaction: motivation for addition of angular momenta as next topic (sec. 3.8)

Summary of radial equation solving:

$$\langle x' | E \ell m \rangle = R_{E\ell}(r) Y_{\ell}^m(\theta, \phi)$$

$R_{E\ell}(r)$ Y_{ℓ}^m eigenstate

requiring "sensible" solution determines energy levels

$R_{E\ell}(r) = U_{E\ell}(r)/r$ makes it "1d"
 Schrodinger wave equation, with
 $V_{\text{eff}}(r) = V(r) + \frac{\ell(\ell+1)\hbar^2}{2mr^2}$ } angular momentum barrier

(1) Short-distance ($r \rightarrow 0$) behaviour:

for $r^2 V(r) \rightarrow 0$ as $r \rightarrow 0$ (e.g., 3d isotropic SHO; 1-electron atom)

\Rightarrow angular momentum barrier (+ sensible solution) dictates $U_{E\ell}(r) \propto r^{\ell+1}$ (for $r \rightarrow 0$)

(2) Long-distance: for $V(r) \rightarrow 0$, thus $V_{\text{eff}} \rightarrow 0$ as $r \rightarrow \infty$, bound states for

$E < 0$... useful: $\rho = \kappa r$, $\kappa = \frac{-2mE}{\hbar^2}$

$\Rightarrow U_{E\ell}(\rho) \propto e^{-\rho}$...

(applies to Coulomb, but **not** to 3d SHO...)

Combining 2 asymptotics,

$U_{E\ell}(\rho) = \underbrace{\rho^{\ell+1}}_{\text{short...}} \times \underbrace{\omega(\rho)}_{\substack{\text{depends} \\ \text{on } V}} \times \underbrace{e^{-\rho}}_{\text{long...}}$

Coulomb potential (also classical gravity)

- One-electron atom with charge of nucleus = $-Ze$ ($e < 0$) $\Rightarrow V = -\frac{Ze^2}{r}$

- goal (general): solve for radial function, getting energy eigenvalues

- Satisfies both $V(r)$ conditions: as $r \rightarrow 0$
& $r^2 V \rightarrow 0$ as $r \rightarrow \infty$ so that

$$u_{El}(\rho) = \rho^{\ell+1} w(\rho) e^{-\rho}$$

- Use $\rho_0 = \sqrt{2m|E|} \left| \frac{Ze^2}{\hbar} \right| = \sqrt{2mc^2 Z\alpha} \sqrt{-E}$

[$\alpha = e^2/(\hbar c) \approx \frac{1}{137}$ fine structure constant]

- General equation for $w(\rho)$ becomes

$$\rho \frac{d^2 w}{d\rho^2} + 2(\ell+1-\rho) \frac{dw}{d\rho} + [\rho_0 - 2(\ell+1)] w = 0$$

- Goal (specifically): for given ℓ , what ρ_0 (or E) will give "reasonable" solution?

- Solve by series ansatz: see sec. 4.2.1 of (3rd edition of) Griffiths

where $\omega \rightarrow \nu$ in Eq. 4.61

- Summary (like 3d isotropic SHO):

- (1). Plug series into DE: set net coefficient of ρ^n in $DE = 0 \dots$
 - (2). recursion relations for coefficients in $\omega \dots$
 - (3). Require normalizability of $\omega(\rho) \dots \Rightarrow$ series must terminate $\dots \Rightarrow$ condition on $\rho_0 \dots$ and $\omega(\rho)$ is polynomial
- Result: $\rho_0 = 2n$ ($n = 0, 1, 2, \dots$)
($l = 0, 1, \dots, (n-1)$) principal quantum number)

i.e.,
$$E = -\frac{1}{2} \frac{m c^2 z^2 \alpha^2}{n^2} = -13.6 \text{ eV} \frac{z^2}{n^2}$$

- $\omega(\rho)$ (polynomial, defined by recursion formula, up to normalization) are (associated) Lauguerre

$$= L_{\substack{2l+1 \\ n-l}}(2\rho), \text{ where (mathematically)}$$

"physics notation"

$$L_q^p(x) \equiv (-1)^p \left(\frac{d}{dx} \right)^p L_{p+q}(x);$$

$$L_q(x) \equiv e^x / q! \left(\frac{d}{dx} \right)^q (e^{-x} x^q)$$

so that (including normalization)

$$\Psi_{nlm}(\vec{x}) \equiv \langle \vec{x} | nlm \rangle = R_{nl}(r) Y_l^m(\theta, \phi),$$

$$R_{nl}(r) = \left(\frac{2z}{na_0} \right)^3 \frac{(n-l-1)!}{2n(n+l)!} \boxed{e^{-zr/(na_0)}}$$

↓
 $e^{-\rho}$
(long...)

$$\times \underbrace{L_{\substack{2l+1 \\ n-l-1}} \left(\frac{2zr}{na_0} \right)}_{w(\rho)} \times \left(\frac{2zr}{na_0} \right)^l \rightarrow \frac{\rho^{l+1}}{\rho} \text{ (short)}$$

where Bohr radius a_0 is natural length scale $= \hbar / (m c \alpha) \approx 0.53 \text{ \AA}$:

$$\rho = \kappa r, \text{ with } \kappa = \sqrt{\frac{-2mE}{\hbar^2}} \text{ (general)}$$

Plug $E = -\frac{1}{2} m c^2 z^2 \alpha^2 / n^2$ to get

$$\frac{1}{\kappa} = \frac{\hbar}{m c \alpha} \frac{n}{z} \equiv a_0 n / z$$

— Equivalently (ala Sakurai),
radial equation for $w(\rho)$ written
as (known) Kummer's equation,
whose solution is (already) known
in series form: require
normalizability ... same ρ_0

[Problem 3.22 of Sakurai: 2
methods directly connected:
(associated) Laguerre polynomials (defined
using generating function) solve Kummer's
equation ...]

Properties of energies/radial functions

for given n [$l = 0, 1, \dots, (n-1)$]

(1). Wavefunctions at origin:

only $l = 0$ is non-vanishing angular
(agrees with general case: momentum barrier)

(2). Number of nodes: $(n-1)$ for $l = 0$
vs. none for $l = n-1$ (highest l)

(3) level of degeneracy (E depends

only on n , not on l, m / number of l values

$$\sum_{l=0}^{n-1} \sum_{m=-l}^{+l} 1 = \sum_{l=0}^{n-1} \underbrace{(2l+1)}_{\text{(usual) } \sum m} = \underbrace{2}_{l=0} \sum_{l=1}^{n-1} l + \underbrace{n}_{\substack{\text{no} \\ l=0 \\ \text{here}}} = 2 \left[\frac{(n-1)n}{2} \right] + n = n^2$$

[use $(1+2+\dots+n) = \frac{n(n+1)}{2}$]

(4) Compare to (old) Bohr model of hydrogen atom: same energy levels
[Balmer formula (empirical)] ... from quantization of angular momentum $= \hbar n$

[1-to-1] correspondence between angular momentum and energy

⇒ Bohr model ground state ($n=1$) has non-zero angular momentum ($l=1$)

vs. Schrodinger/Heisenberg theory:

$n=1$ has $l=0$ (zero angular momentum)

and for higher n , $l=0, 1, \dots, (n-1) \dots$

(range of l giving same E)

... degeneracy in l "broken" by

spin-orbit interaction: so far only

energy/angular momentum from (only)

spatial/orbital motion included
(magnetic moment)

- Enter spin of electron (for now!):

contributes to energy (Hamiltonian)

if there is magnetic field (\vec{B})

- \vec{B} acting on electron in one-electron atom ?!

... Yes! Go to **rest** frame of electron,
nucleus orbiting electron \Rightarrow current..
 $\Rightarrow \vec{B}$ acting on electron $\propto \vec{L}$ of proton/
electron
 \Rightarrow contribution to $H \propto \vec{B} \cdot \vec{S}$
 $\propto \boxed{\vec{L} \cdot \vec{S}}$ (fine structure: E
depends on l ; chapter 5 in Phys 623)
 \Rightarrow motivates "adding" angular
momenta