## Problems

## 1. SMM, Chapter 2, Problem 2.

Assume that your skin can be considered a blackbody (which is probably not a very good assumption). One can then use Wien's displacement law,
$\lambda_{\text {max }}(m) T(K)=2.898 \times 10^{-3} m \cdot K$, to find the wavelength of the peak emission.
Letting $T=35$ Celsius $=308 \mathrm{~K}$,

$$
\lambda_{\max }=2.898 \times 10^{-3} \mathrm{~m} \cdot \mathrm{~K} / 308 \mathrm{~K}=9.41 \times 10^{-6} \mathrm{~m}=9409 \mathrm{~nm}
$$

which is in the far infrared (heat).

## 2. SMM, Chapter 2, Problem 3.

According to classical mechanics, for a simple harmonic oscillator having amplitude $A$ and spring constant $k$...

$$
E_{\text {total }}=\frac{1}{2} k A^{2}, f=\frac{1}{2 \pi} \sqrt{\frac{k}{m}}
$$

(a) $E_{\text {total }}=\frac{1}{2}(25 \mathrm{~N} / \mathrm{m})(0.4 \mathrm{~m})^{2}=2.0 \mathrm{~J}$, and $f=\frac{1}{2 \pi} \sqrt{\frac{(25 \mathrm{~N} / \mathrm{m})}{(2 \mathrm{~kg})}}=\mathbf{0 . 5 6 ~ H z}$.
(b) If the energy is quantized, it will be given by $E_{n}=n h f$. We find the value of the quantum number $n$ simply by,

$$
n=\frac{E_{n}}{h f}=\frac{(2.0 J)}{\left(6.63 \times 10^{-34} J \cdot s\right)(0.56 \mathrm{~Hz})}=5.4 \times 10^{33} .
$$

(c) At these frequencies, the energy carried away by one quantum change will be,

$$
E=h f=\left(6.63 \times 10^{-34} J \cdot s\right)(0.56 \mathrm{~Hz})=3.7 \times 10^{-34} \mathrm{~J}
$$

## 3. SMM, Chapter 2, Problem 4.

(a) Stefan's law tells us that the power per unit area emitted at all frequencies by a blackbody is proportional to the fourth power of its absolute temperature. Thus,

$$
e_{\text {total }}=\sigma T^{4}=\left(5.67 \times 10^{-8} \mathrm{Wm}^{-2} K^{-4}\right)(3000 \mathrm{~K})^{4}=4.59 \times 10^{6} \mathbf{W} / \mathbf{m}^{2}
$$

(b) If a light bulb is rated for $\mathrm{P}=75 \mathrm{~W}$, then...

$$
e_{\text {total }}=\frac{\text { Power }}{\text { Area }} \rightarrow \text { Area }=\frac{\text { Power }}{e_{\text {total }}}=\frac{(75 \mathrm{~W})}{\left(4.59 \times 10^{6} \mathrm{~W} / \mathrm{m}^{2}\right)}=\mathbf{1 6 . 3} \mathbf{~ m m}^{2} .
$$

4. Calculate the energy of a photon whose frequency is (a) $5 \times 10^{14} \mathrm{~Hz}$, (b) 10 GHz , (c) 30 MHz . Express your answers in electron volts. Also determine the
corresponding wavelengths for each case and what part of the EM spectrum this is. (This is essentially SMM problem $2.7 \& 2.8$ combined).

In all cases we use $E_{n}=n h f$ and $\lambda=\frac{c}{f} \ldots$
(a) $E=\left(6.63 \times 10^{-34} \mathrm{~J} . \mathrm{s}\right)\left(5 \times 10^{14} \mathrm{~Hz}\right)=3.315 \times 10^{-19} \mathrm{~J}=2.07 \mathrm{eV}$
$\lambda=\left(3 \times 10^{8} \mathrm{~m} / \mathrm{s}\right) /\left(5 \times 10^{14} \mathrm{~Hz}\right)=6 \times 10^{-7} \mathrm{~m}=\mathbf{6 0 0} \mathrm{nm}$, visible (yellow)
(b) $E=\left(6.63 \times 10^{-34} \mathrm{~J} . \mathrm{s}\right)\left(10 \times 10^{9} \mathrm{~Hz}\right)=6.63 \times 10^{-24} \mathrm{~J}=4.14 \times 10^{-5} \mathrm{eV}$

$$
\lambda=\left(3 \times 10^{8} \mathrm{~m} / \mathrm{s}\right) /\left(10 \times 10^{9} \mathrm{~Hz}\right)=\mathbf{0 . 0 3} \mathbf{~ m} \text {, microwave }
$$

(c) $E=\left(6.63 \times 10^{-34} \mathrm{~J} . \mathrm{S}\right)\left(30 \times 10^{6} \mathrm{~Hz}\right)=1.989 \times 10^{-26} \mathrm{~J}=1.24 \times 10^{-7} \mathrm{eV}$

$$
\lambda=\left(3 \times 10^{8} \mathrm{~m} / \mathrm{s}\right) /\left(30 \times 10^{6} \mathrm{~Hz}\right)=10 \mathrm{~m}, \text { radio }
$$

## 5. SMM, Chapter 2, Problem 11.

In general, Power $=\frac{\text { energy }}{\text { time }}=\frac{n h f}{t}=\left(\frac{n}{t}\right) h f$. Isolating for the number of photons per unit time...
$\frac{n}{t}=\frac{\text { photons }}{\text { time }}=\frac{\text { Power }}{h f}=\frac{P}{h}\left(\frac{\lambda}{c}\right)=\frac{(10 \mathrm{~W})\left(589.3 \times 10^{-9} \mathrm{~m}\right)}{\left(6.63 \times 10^{-34} \mathrm{~J} \cdot \mathrm{~s}\right)\left(3 \times 10^{8} \mathrm{~m} / \mathrm{s}\right)}=2.96 \times 10^{19}$ photons $/ \mathbf{s}$.
6. Extra Credit problem 1. Using Planck's spectral distribution formula, $u(\lambda, T)$, and recalling that $e_{\text {total }}=\frac{c}{4} \int_{0}^{\infty} u(\lambda, T) d \lambda$, derive Stefan's law, $e_{\text {total }}=\sigma T^{4}$, for the total power per unit area radiated at all wavelengths. Work out the numerical value for the constant $\sigma$. Useful hint: $\int_{0}^{\infty} \frac{x^{3}}{\left(e^{x}-1\right)} d x=\frac{\pi^{4}}{15}$.

This problem is worked out in complete detail on page 70 of SMM. The derivation goes something like this.

$$
e_{\text {total }}=\frac{c}{4} \int_{0}^{\infty} u(\lambda, T) d \lambda=\frac{c}{4} \int_{\lambda=0}^{\lambda=\infty} \frac{8 \pi h c}{\lambda^{5}\left(e^{h c / \lambda k_{b} T}-1\right)} d \lambda=\frac{2 \pi k_{b}^{4} T^{4}}{h^{3} c^{2}} \int_{x=0}^{x=\infty} \frac{x^{3}}{\left(e^{x}-1\right)} d x
$$

where we make use of the change of variable $x \equiv h c / \lambda k_{b} T$. Note: In making the change of variables, don't forget to substitute $d \lambda=-\frac{h c}{x^{2} k_{b} T} d x$, and the limits of integration. There's a little bit of algebra in doing this.

Finally, using the value for the integral...

$$
e_{\text {total }}=\frac{2 \pi^{5} k_{b}^{4}}{15 c^{2} h^{3}} T^{4}=\sigma T^{4}
$$

Substituting for $k_{b}, c$, and $h$, we have

$$
\sigma=\frac{2 \pi^{5}\left(1.381 \times 10^{-23} \mathrm{~J} / \mathrm{K}\right)^{4}}{15\left(3 \times 10^{8} \mathrm{~m} / \mathrm{s}\right)^{2}\left(6.626 \times 10^{-34} \mathrm{~J} \cdot \mathrm{~s}\right)^{3}}=5.67 \times 10^{-8} \mathrm{~W} \cdot \mathrm{~m}^{-2} \cdot \mathrm{~K}^{-4}
$$

7. Extra Credit problem 2. Using Planck's spectral distribution formula, $u(\lambda, T)$, (a) Derive Wein's displacement law, $\lambda_{\max } T=$ const. Assume that the transcendental equation, $x=5\left(1-e^{-x}\right)$, has a non-trivial solution given by $\mathrm{x}_{0}$. This comment will be clear as you work through the problem.
(b) Using a dimensionless value for the non-trivial solution of $x_{0}=4.96511423$, work out the value and units of the constant.

In principle, the procedure is simple. To find the peak wavelength of the spectral distribution formula, we take the derivative with respect to wavelength, set it to zero and solve for the wavelength. And so ...

$$
u(\lambda, T)=\frac{8 \pi h c}{\lambda^{5}\left(e^{h c / \lambda k_{b} T}-1\right)}
$$

Setting the derivative to zero,

$$
\frac{\partial u}{\partial \lambda}=8 \pi h c\left[\left(\frac{h c}{k_{b} T}\right) \frac{1}{\lambda^{7}} e^{h c / \lambda k_{b} T}\left(\frac{1}{e^{h c / \lambda k_{b} T}-1}\right)^{2}-\frac{5}{\lambda^{6}}\left(\frac{1}{e^{h c / \lambda k_{b} T}-1}\right)\right]=0
$$

Solving for $\lambda_{\max } \ldots$ (Note: All the $\lambda$ 's in the equations are $\lambda_{\max }$ but for brevity, I'm just calling them $\lambda$.)

$$
\begin{gathered}
\left(\frac{h c}{k_{b} T}\right) \frac{1}{\lambda^{7}} e^{h c / \lambda k_{b} T}\left(\frac{1}{e^{h c / \lambda k_{b} T}-1}\right)^{2}=\frac{5}{\lambda^{6}}\left(\frac{1}{e^{h c / \lambda k_{b} T}-1}\right) \\
\frac{1}{5}\left(\frac{h c}{\lambda k_{b} T}\right) e^{h c / \lambda k_{b} T}=e^{h c / \lambda k_{b} T}-1 \\
\frac{h c}{\lambda k_{b} T}=5\left(1-e^{-h c / \lambda k_{b} T}\right)
\end{gathered}
$$

Letting $x \equiv h c / \lambda k_{b} T$ we have

$$
x=5\left(1-e^{-x}\right)
$$

If we want to solve for $\lambda_{\max }$ we must solve for $x$. This is a transcendental equation in $x$ that will not have an analytical solution but a numerical one. For the time being, let's assume that this numerical solution is $x_{0}$. And so,

$$
\begin{gathered}
x=x_{0} \\
\frac{h c}{\lambda_{\max } k_{b} T}=x_{0} \\
\lambda_{\max } T=\frac{h c}{x_{0} k_{b}}=\text { const. }
\end{gathered}
$$

Assuming that $x_{0}=4.96511423$ (which is dimensionless), we can solve for the constant.

$$
\lambda_{\max } T=\frac{\left(6.626 \times 10^{-34} \mathrm{~J} \cdot \mathrm{~s}\right)\left(3 \times 10^{8} \mathrm{~m} / \mathrm{s}\right)}{(4.96511423)\left(1.381 \times 10^{-23} \mathrm{~J} / \mathrm{K}\right)}=0.002899 \mathrm{~m} \cdot \mathrm{~K}
$$

