February 15, $2017 \quad$ Physics $132 \quad$ Prof. E. F. Redish
$■$ Theme Music: Zimmer \& Howard
Agents of Chaos
■ Cartoon: Bob Thaves
Frank \& Ernest
Frank and Ernest


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## Quiz 3

|  | $\mathbf{1}$ | $\mathbf{2}$ | $\mathbf{3}$ | $\mathbf{4}$ |
| :---: | :---: | :---: | :---: | :---: |
| A | $11 \%$ | $2 \%$ | $24 \%$ | $16 \%$ |
| B | $2 \%$ | $6 \%$ | $42 \%$ | $85 \%$ |
| C | $82 \%$ | $1 \%$ | $2 \%$ | $0 \%$ |
| D | $5 \%$ | $1 \%$ | $32 \%$ |  |
|  |  | $86 \%$ | $1 \%$ |  |
|  |  |  |  |  |

## Quiz 3




## Foothold ideas: <br> The Second Law of Thermodynamics

Systems spontaneously move toward the thermodynamic (macro)state that correspond to the largest possible number of particle arrangements (microstates).

- The $2^{\text {nd }}$ law is probabilistic. Systems show fluctuations violations that get proportionately smaller as N gets large.
- Systems that are not in thermodynamic equilibrium will spontaneously transform so as to increase the entropy.
- The entropy of any particular system can decrease as long as the entropy of the rest of the universe increases more.
■ The universe tends towards states of increasing chaos and uniformity. (Is this contradictory?)


## Foothold ideas: <br> Energy distribution

■ Due to the randomness of thermal collisions, even in (local) thermal equilibrium the energy in each DoF fluctuates, so a range of energy will be found in each degree of freedom.
$\square$ The probability of adding an energy $\Delta E$ is proportional to the Boltzmann factor

$$
\begin{aligned}
& P(\Delta E) \propto e^{-\Delta E / k_{B} T}(\text { for one DoF) } \\
& P(\Delta E) \propto e^{-\Delta E / R T}(\text { for one mole })
\end{aligned}
$$

■ At 300 K ,

$$
\begin{aligned}
& k_{\mathrm{B}} T \sim 1 / 40 \mathrm{eV}=25 \mathrm{meV} \\
& N_{\mathrm{A}} k_{\mathrm{B}} T=R T \sim 2.4 \mathrm{~kJ} / \mathrm{mol}
\end{aligned}
$$

## The Boltzmann probability

- The probability of finding an additional energy $\Delta E$ in a DoF is proportional to the number of ways that that energy can be distributed, $W$.
- The overall probability has to be normalized so that the sum (integral) over all energies is 1 .

$$
P(\Delta E, T)=P_{0} W(\underbrace{\Delta E, T)} e^{-\Delta E / k_{B} T}
$$

The number of ways $\Delta E$ can be distributed at a temperature $T$

