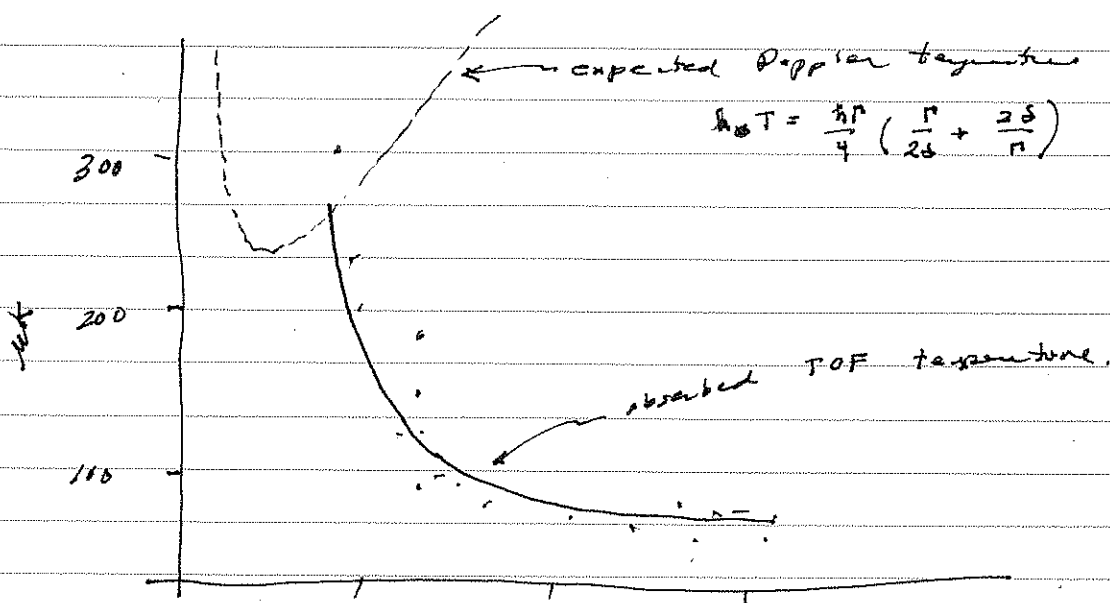


8 December 2005 Sub-Doppler laser cooling continued.

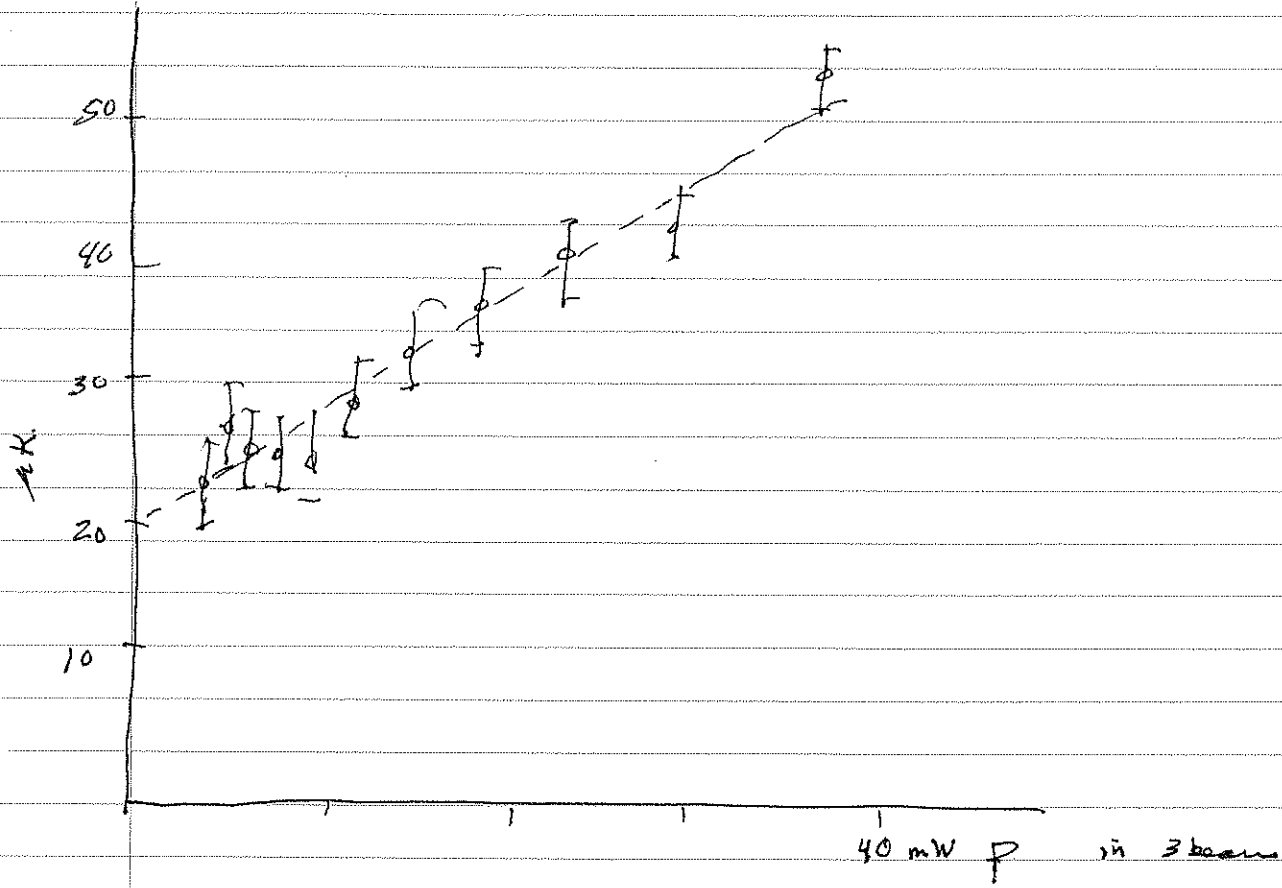
The sub-Doppler temperature seen with time-of-flight was confirmed by three other methods

The temperature depended on detuning in a way different from predicted by Doppler cooling



Also: the temperature went up dramatically when a small magnetic field was applied, even when

$$\Delta W_{Zeeman} \ll \Gamma$$



the temperature was linearly dependent on the intensity.

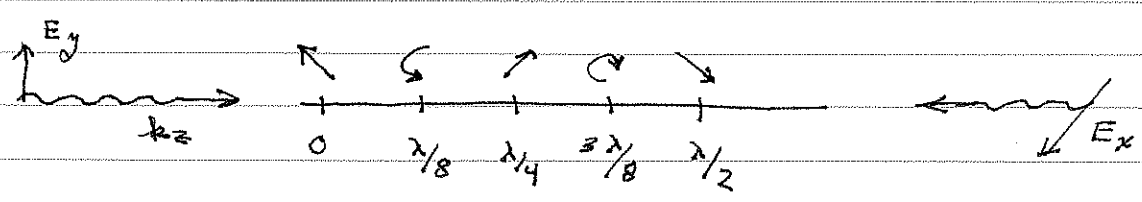
What is going on? Doppler cooling theory is simple! How can it be wrong?

hint: magnetic field dependence suggest magnetic sub levels are involved.

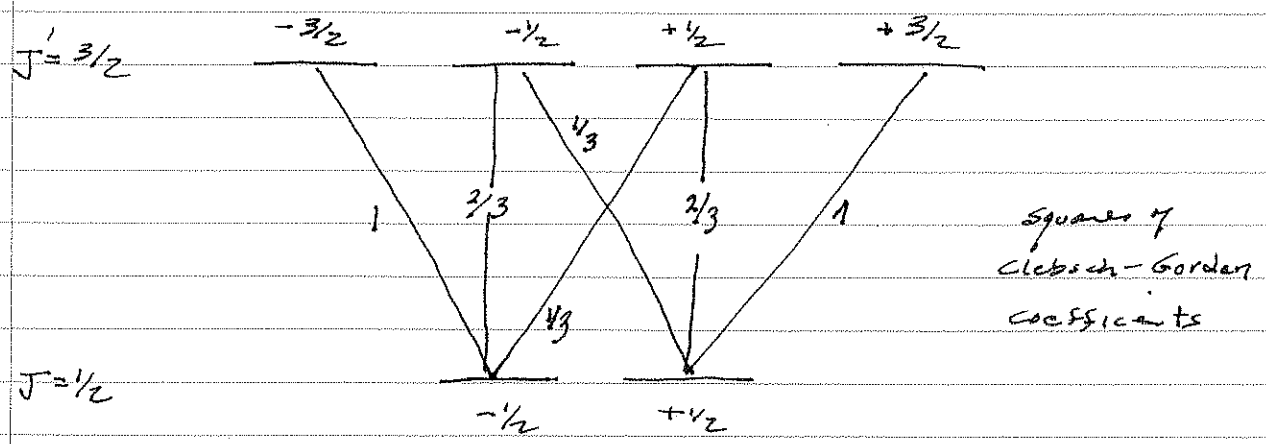
The answer (from Paris - ENS) involved:

- multi-level atoms
- optical pumping among the sub-levels.
- differential light shifts
- polarization gradient of the light.

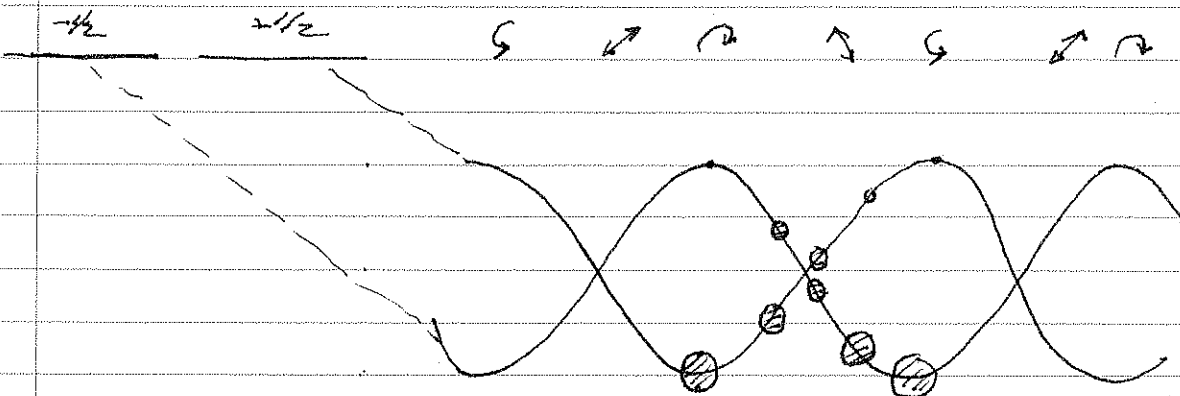
polarization gradient



transition strengths

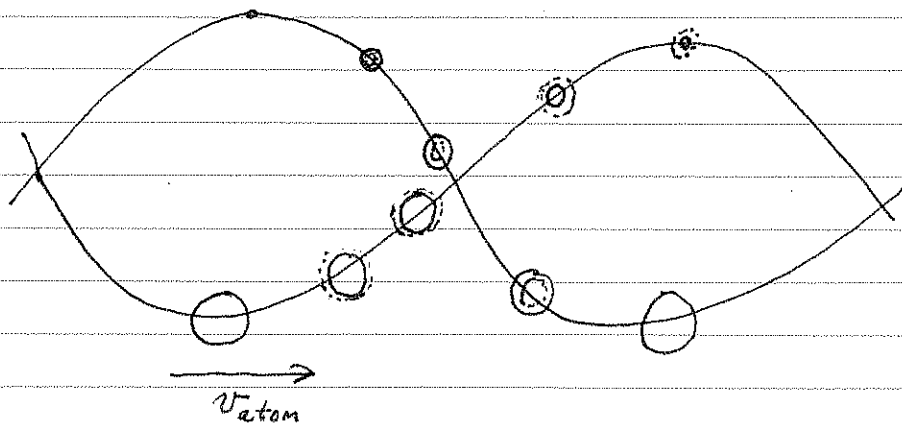


light shifts



atoms pump to the lowest level (when  $\delta < 0$  for negative light shift) - populations for static atoms.

When atoms move, populations lag because pumping takes time - longer time for low saturation and large detuning

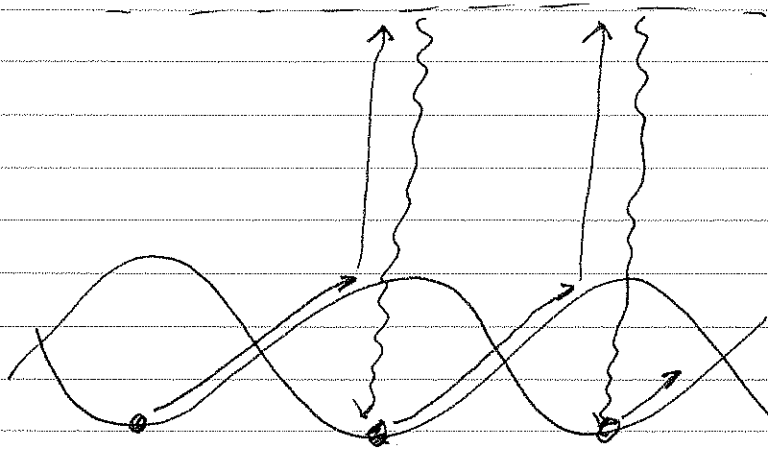


population is larger going up

smaller going down

net result is slowing of atom

Simplified picture:

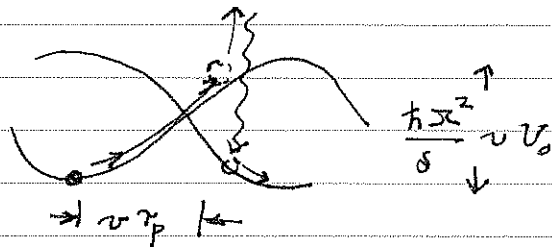


atom pumps after  $\tau_p$ , where

$$\tau_p^{-1} \sim \frac{\Gamma}{2} \frac{I/I_0}{1 + (2\delta/\Gamma)^2} \quad \text{for } I/I_0 \ll 1$$

$$\text{or } \tau_p \sim \frac{\Gamma \frac{\Omega^2/\Gamma^2}{\delta^2/\Gamma^2}}{\delta^2/\Gamma^2} \sim \frac{\Gamma \Omega^2}{\delta^2} \quad \text{for } \delta \gg \Gamma, \quad I/I_0 \ll 1$$

handwaving calculation of average force:



$$\langle F \rangle \sim U_0 \left( \frac{\nu \tau_p}{\lambda} \right) \frac{1}{\lambda} \quad \text{or} \quad \frac{\hbar \Omega^2 k^2 \nu}{\delta \Gamma \Omega^2 / \delta^2} = \frac{3 \hbar k^2 \delta}{\Gamma} \nu$$

total potential

fraction of potential used

characteristic distance over which potential changes

compare to

$$F_{\text{dop}}^{\text{max}} = \frac{\hbar k^2}{4} \nu$$

Force is independent of  $\Omega^2$ , and much bigger than  $F_{Dopp}$  for  $\delta \gg \Gamma$ .

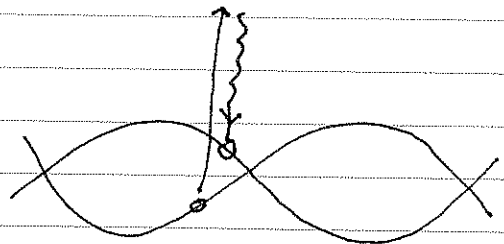
This larger friction explains the odd behavior of molasses lifetime.

What about temperature? What about  $\Omega^2 \rightarrow 0$ ?

what is the heating? - momentum diffusion.

recall.  $\hbar \Gamma = \frac{d\langle p \rangle}{dt}$

$\Omega \frac{d\langle p \rangle}{dt} = \langle p^2 \rangle$



$\Omega \Gamma \sim \underbrace{(F \cdot r_p)^2}_{\text{impulse on one potential}} \underbrace{\frac{1}{\Gamma_p}}_{\text{rate of impulses}}$

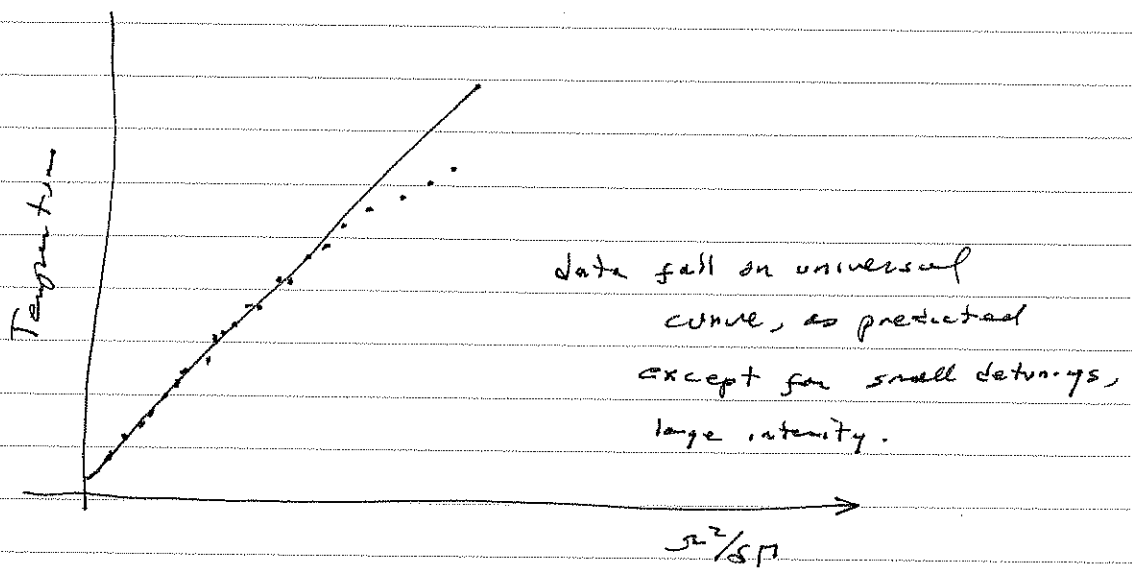
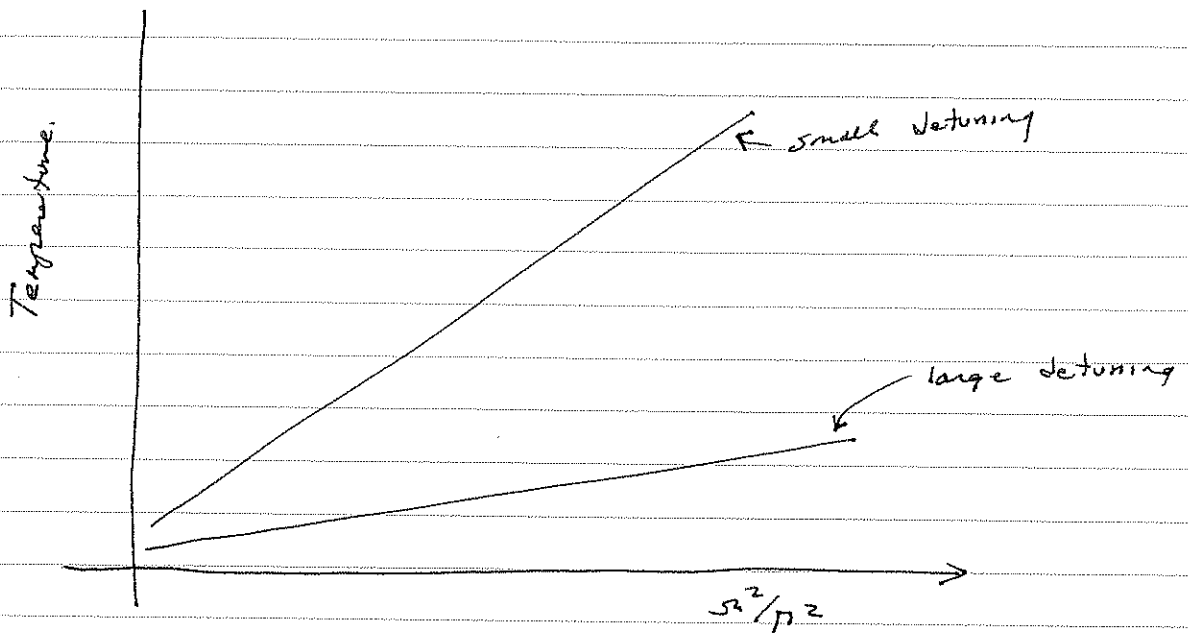
force fluctuates as atom pumps between potentials.

$\Omega \Gamma \sim \left( \frac{\hbar \Omega^2}{\delta} k \right)^2 \frac{1}{\Gamma \frac{\Omega^2}{\delta^2}} \sim \frac{\hbar^2 k^2 \Omega^2}{\Gamma}$

$\hbar \Omega_0$  derivative characteristic length

depends on intensity

$\hbar \Gamma = \frac{d\langle p \rangle}{dt} \sim \frac{\hbar^2 k^2 \Omega^2}{\Gamma \hbar^2 \delta / \mu} \sim \frac{\hbar \Omega^2}{\delta} \quad (\sim \text{light shift})$



What about  $\Omega^2 \rightarrow 0$ ? Does  $T \rightarrow 0$ ?

No. (obviously)

the calculated  $\Delta p \sim \Omega^2$  is only the part due to fluctuations of the dipole force

there is another part, small for large  $\Omega^2$ , due to photon recoil.

where  $U_0 \sim E_{\text{recoil}}$ , the atom loses less energy than it gains in an optical pump cycle, so cooling stops.

$$kT_{\text{min}} \sim E_{\text{rec}}$$

In practice  $kT_{\text{min}} \approx 4E_{\text{rec}}$  ~~was~~ measured for Cs.

$$\text{but } \frac{E_{\text{rec}}}{k} = 200\text{K for Cs}$$

$$\text{while } T_{\text{Dop}}^{\text{min}} \approx 120\text{K for Cs.}$$

so, "polarization-gradient" or "Sisyphus" cooling is

much more effective than Doppler cooling.

This enabled things like fountain clocks and Bose-Einstein  
Condensates.

short Topics :

Ramsey method

atomic clocks

fountain clocks

magnetic trapping