

Discussion of photons brings us to quantized light:

How do we move to quantized fields?

replace

$$\vec{E} \rightarrow \hat{\vec{E}} = \sum_{\mathbf{k}, \lambda} i \epsilon_{\mathbf{k}\lambda} \hat{\epsilon}_{\mathbf{k}\lambda} (\hat{a}_{\mathbf{k}} e^{i\mathbf{k}\cdot\vec{r}} + \hat{a}_{\mathbf{k}}^{\dagger} e^{-i\mathbf{k}\cdot\vec{r}})$$

note: some authors subsume  $i$  into  $\hat{a}$ ,  $(i\hat{a})^{\dagger} = -i\hat{a}^{\dagger}$   $i\hat{a}^{\dagger} = -i\hat{a}$   
can also be written

$$\hat{\vec{E}} = \sum_{\mathbf{k}, \lambda} \epsilon_{\mathbf{k}\lambda} \hat{\epsilon}_{\mathbf{k}\lambda} (\hat{a}_{\mathbf{k}} e^{i\mathbf{k}\cdot\vec{r}} + \hat{a}_{\mathbf{k}}^{\dagger} e^{-i\mathbf{k}\cdot\vec{r}})$$

Problem: what is the intensity?

$$\vec{E}^{(+) }(\vec{r}) = \sum_{\mathbf{k}, \lambda} \epsilon_{\mathbf{k}\lambda} \hat{\epsilon}_{\mathbf{k}\lambda} \hat{a}_{\mathbf{k}} e^{i\mathbf{k}\cdot\vec{r}}$$

$$\vec{E}^{(-) }(\vec{r}) = \sum_{\mathbf{k}, \lambda} \hat{\epsilon}_{\mathbf{k}\lambda} \hat{a}_{\mathbf{k}}^{\dagger} e^{-i\mathbf{k}\cdot\vec{r}}$$

$$\frac{1}{2} \epsilon_0 c |\vec{E}|^2 = c U = c \sum_{\mathbf{k}} \hbar \omega_{\mathbf{k}} (\hat{a}_{\mathbf{k}}^{\dagger} \hat{a}_{\mathbf{k}} + \frac{1}{2})$$

↑ energy density diverges.

Look at detection from a quantum point of view:  
use photo-excitation/ionization as a model.

starting in state  $|i\rangle = |n_{\mathbf{k}}, \Phi_0\rangle$ , where  $|\Phi_0\rangle$  is the state of the detector, we detect light when photons make transitions to  $|\Phi_s\rangle$

e.g.

$$\langle \Phi_s | \hat{\vec{E}} | \Phi_0 \rangle \rightarrow \text{matrix element.}$$

in any event, the radiative part enters independently of the atomic (detector part) and ~~with~~ its contribution to the rate is (for a single mode)

$$|\langle n_{\mathbf{k}} - 1 | \hat{\vec{E}} | n_{\mathbf{k}} \rangle|^2$$

Only the destruction operator part contributes,

$$\langle n_k - 1 | \hat{E}^+(r) | n_k \rangle^2 = \langle n_k | \hat{E}^-(r) | n_k - 1 \rangle \langle n_k - 1 | \hat{E}^+(r) | n_k \rangle$$

~~is~~  $\hat{E}$  is not

a vector, its the  
magnitude.

$$= \langle n_k | \hat{E}^{(-)}(r) \hat{E}^{(+)}(r) | n_k \rangle.$$

→ called normally ordered.

For a statistical average, ~~this part is~~

~~The~~ The photon contribution to the rate is

$$\sum_{n_k} P(n_k) \langle n_k | \hat{E}^{(-)}(r) \hat{E}^{(+)}(r) | n_k \rangle = \text{Tr}(\hat{\rho} \hat{E}^-(r) \hat{E}^+(r))$$

(the atomic part drops out as well, if you sum over all possible final detector states leaving you with an overall quantum efficiency)

this can be generalized to a quantum-optical Poynting vector

$$\hat{\mathbf{I}} = \epsilon_0 c^2 \left( \hat{\mathbf{E}}^{(-)}(r) \times \hat{\mathbf{B}}^{(+)}(r) - \hat{\mathbf{B}}^{(-)}(r) \times \hat{\mathbf{E}}^{(+)}(r) \right)$$

which for polarized parallel light gives

$$\hat{\mathbf{I}} = 2\epsilon_0 c \hat{E}^{(-)}(r) \hat{E}^{(+)}(r) \frac{\vec{k}}{|\vec{k}|}$$

and

$$\langle \hat{\mathbf{I}} \rangle = \frac{c}{V} \sum_k \hbar \omega_k \langle n_k \rangle = \frac{c}{V} (\langle E_R \rangle - E_{vac})$$

Use Heisenberg operators

$$g^{(1)}(\vec{r}_1, \vec{r}_2, \tau) = \frac{\langle \hat{E}^{(-)}(\vec{r}_1, t) \hat{E}^{(+)}(\vec{r}_2, t+\tau) \rangle}{\sqrt{\langle \hat{E}^{(-)}(\vec{r}_1) \hat{E}^{(+)}(\vec{r}_1) \rangle \langle \hat{E}^{(-)}(\vec{r}_2) \hat{E}^{(+)}(\vec{r}_2) \rangle}}$$

$$g^{(2)}(\vec{r}_1, \vec{r}_2, \tau) = \frac{\langle \hat{E}^{(-)}(\vec{r}_1, t) \hat{E}^{(-)}(\vec{r}_2, t+\tau) \hat{E}^{(+)}(\vec{r}_2, t+\tau) \hat{E}^{(+)}(\vec{r}_1, t) \rangle}{\langle \hat{E}^{(-)}(\vec{r}_1) \hat{E}^{(+)}(\vec{r}_1) \rangle \langle \hat{E}^{(-)}(\vec{r}_2) \hat{E}^{(+)}(\vec{r}_2) \rangle}$$

When  $\langle \hat{O} \rangle = \text{Tr}(\hat{\rho} \hat{O})$

$$\langle I_{\text{MZ}} \rangle = \frac{1}{2} \epsilon_0 \frac{1}{4} \langle (\hat{E}^{(-)}(t_1) + \hat{E}^{(-)}(t_2)) (\hat{E}^{(+)}(t_1) + \hat{E}^{(+)}(t_2)) \rangle$$

One can show  $0 \leq g^{(2)}(\tau) \leq \infty$ , but  
one cannot show that  $g^{(2)}(\tau) \geq 1$ !

Examples  $\rightarrow$  start with single mode only, for examples  
 $\rightarrow$  note: single mode is monochromatic,  
classically we saw that

$$F(\omega) = \frac{1}{2\pi} \int_{-\infty}^{\infty} d\tau g^{(1)}(\tau) e^{i\omega\tau}$$

$\rightarrow g^{(1)}(\tau) \rightarrow \text{constant}$ .

$\rightarrow$  light is a sine-wave of fixed phase,  $\phi$   
pretty boring deterministic light.

$\phi$  is the only variable freedom

Scale electric field operator so that

$$\hat{E}(\phi) = \hat{E}^+(\phi) + \hat{E}^-(\phi) = \frac{1}{2} \hat{a} e^{-i\phi} + \frac{1}{2} \hat{a}^\dagger e^{i\phi}$$

$$\rightarrow = \hat{X} \cos \phi + \hat{Y} \sin \phi$$

quadrature expression

$$\phi = \omega t - kx + \text{const}$$

$$\bullet \left[ \hat{E}(\phi_1), \hat{E}(\phi_2) \right] = -\frac{i}{2} \sin(\phi_1 - \phi_2)$$

$$\hat{X} = \frac{\hat{a} + \hat{a}^\dagger}{2} \quad \hat{Y} = \frac{\hat{a} - \hat{a}^\dagger}{2i}$$

$\hookrightarrow \hat{E}$  at different  $\phi$  don't commute.

For single-mode

$$g^{(1)}(x_1, x_2, \tau) = \frac{\langle \frac{1}{2} \hat{a} e^{i\phi_2} \frac{1}{2} \hat{a} e^{-i\phi_1} \rangle}{\sqrt{\langle \frac{1}{4} \hat{a}^\dagger \hat{a} \rangle \langle \frac{1}{4} \hat{a}^\dagger \hat{a} \rangle}}$$

$$= e^{i(\phi_2 - \phi_1)}$$

$$\phi_1 = -kx_1 + \omega t$$

$$\phi_2 = -kx_2 + \omega(t + \tau)$$

$$|g^{(1)}(\tau)| = 1 \rightarrow \text{kind of expected.}$$

$$\rightarrow \text{define } S(\hat{E}(\phi)) = \langle \hat{E}(\phi) \rangle$$

$$N(E(\phi)) = (\Delta E(\phi))^2 \quad (\Delta E(\phi))^2 = \langle \hat{E}^2 \rangle - \langle \hat{E} \rangle^2$$

What about  $g^{(2)}(\tau)$ ?

$$g^{(2)}(x_1, x_2, \tau) = \frac{1}{\bar{n}^2} \langle \hat{a}^\dagger e^{i\phi_2} \hat{a}^\dagger e^{i\phi_2} \hat{a} e^{-i\phi_1} \hat{a} e^{-i\phi_1} \rangle$$

$$\frac{\langle \hat{a}^\dagger \hat{a} \rangle \langle \hat{a}^\dagger \hat{a} \rangle}{\bar{n}^2}$$

$$= \frac{\langle \hat{a}^\dagger \hat{a}^\dagger \hat{a} \hat{a} \rangle}{\langle \hat{a}^\dagger \hat{a} \rangle^2}$$

Note:  $g^{(2)}$  doesn't depend on  $\phi$ , (or  $\Delta x$  or  $\tau$ ),

but it does depend on the state that is chosen, and it is not always given by deterministic light!

$$a a^\dagger = a^\dagger a + 1$$

$$g^{(2)}(\tau) = \frac{\langle n(n-1) \rangle}{\langle n \rangle^2} = \frac{\langle n^2 \rangle - \langle n \rangle}{\langle n \rangle^2}$$

add to  
subtract 1

$$= 1 + \frac{\langle \Delta n \rangle^2 - \langle n \rangle}{\langle n \rangle^2}$$

Since  $\langle \Delta n \rangle^2 \geq 0$

$$\langle \Delta n \rangle^2 = \langle n^2 \rangle - \langle n \rangle^2$$

$$g^{(2)}(\tau) \geq 1 - \frac{1}{\langle n \rangle}$$

$g^{(2)}(\tau) \geq 0$  no matter what.

For finite photon number, this is less than 1!  
Statistically impossible for classical light.  
Compared to EPR!

Specific examples:

number or Fock state:

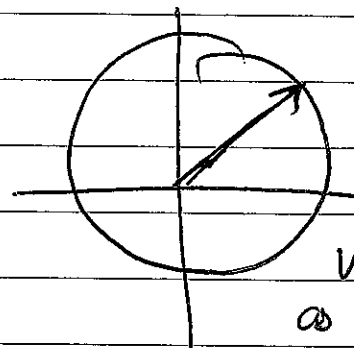
$$(\Delta n)^2 = 0, \quad g^{(2)} = 1 - \frac{1}{n} \text{ for } n \geq 1$$

We showed early in the class, that the quadrature expectation values on number states vanish:

$$\langle n | \hat{X} | n \rangle = \langle n | \hat{Y} | n \rangle = 0$$

but

$$(\Delta X)^2 = (\Delta Y)^2 = \frac{1}{2}(n + \frac{1}{2})$$



viewed  
as all phases  
available.

Coherent states:

$$|\alpha\rangle = e^{-\frac{1}{2}|\alpha|^2} \sum_{n=0}^{\infty} \frac{\alpha^n}{\sqrt{n!}} |n\rangle$$

$$\hat{a} |\alpha\rangle = \alpha |\alpha\rangle$$

$$\langle \alpha | \hat{a}^\dagger = \langle \alpha | \alpha^*$$

$$\langle \alpha | \beta \rangle \neq 0 = e^{-|\alpha - \beta|^2} \rightarrow \text{overcomplete.}$$

$$\langle n \rangle = \langle \alpha | \hat{n} | \alpha \rangle = \langle \alpha | \hat{a}^\dagger \hat{a} | \alpha \rangle = |\alpha|^2$$

$$\begin{aligned} \langle n^2 \rangle &= \langle \alpha | \hat{a}^\dagger \hat{a} \hat{a}^\dagger \hat{a} | \alpha \rangle = \langle \alpha | \hat{a}^\dagger \hat{a} \hat{a}^\dagger \hat{a} + \hat{a}^\dagger \hat{a} | \alpha \rangle \\ &= |\alpha|^4 + |\alpha|^2 = \langle n \rangle^2 + \langle n \rangle \end{aligned}$$

$$\langle \Delta n \rangle^2 = \langle n \rangle$$

→ not like a ~~st~~ deterministic classical wave.

$$\frac{\Delta n}{\langle n \rangle} = \frac{1}{\sqrt{\langle n \rangle}} \quad \text{relative uncertainty,}$$

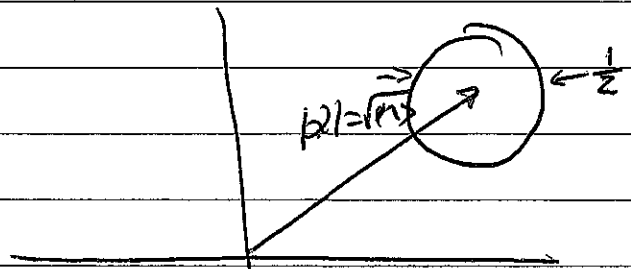
not surprising  $P(n)$  is poissonian

$$|\langle n | \alpha \rangle|^2 = e^{-\langle n \rangle} \frac{\langle n \rangle^n}{n!}$$

Nonetheless,

$$g^{(2)}(\tau) = \frac{\langle \alpha | \hat{a}^{\dagger+\dagger+\dagger} \hat{a}^{\dagger} \hat{a} \hat{a} | \alpha \rangle}{(\langle \alpha | \hat{a}^{\dagger} \hat{a} | \alpha \rangle)^2} = 1 + \frac{\overbrace{(\Delta n)^2 - \langle n \rangle}^{\rightarrow \Delta n^2}}{\underbrace{\langle n \rangle^2}_{=0}}$$

$$g^{(2)}(\tau) = 1 \rightarrow \text{no "bunching"}$$



Chaotic light (single mode)

Thermal distribution of photon number  
(earlier in class)

$$P(n) = \frac{\langle n \rangle^n}{(1 + \langle n \rangle)^{n+1}}$$

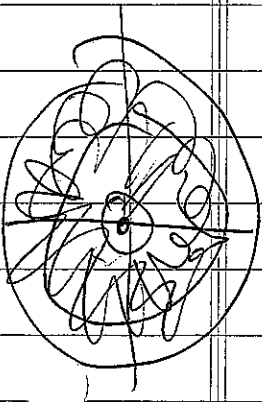
$$(\Delta n)^2 = \langle n \rangle^2 + \langle n \rangle$$

$$g^{(2)}(\tau) = 2 \quad \text{for } \langle n \rangle \text{ large.}$$

definitions:

if  $(\Delta n)^2 - \langle n \rangle > 0 \rightarrow$  super poissonian

$(\Delta n)^2 - \langle n \rangle < 0$  sub poissonian



Quadrature.  
Squeezed light:

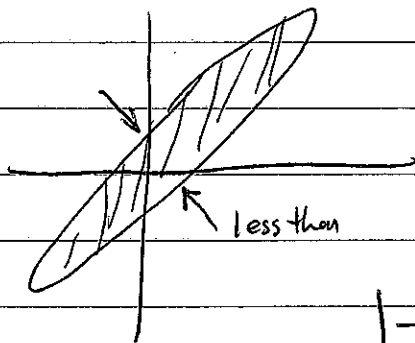
Consider the state  $|\xi\rangle = \hat{S}(\xi)|0\rangle \rightarrow$  squeezed vacuum

$$\hat{S}(\xi) = e^{\left(\frac{1}{2} s^* \hat{a}^2 - \frac{1}{2} s (\hat{a}^\dagger)^2\right)}, \quad s = |s| e^{i\theta}$$

$\rightarrow$  looks like a coherent state, but  $\hat{a}, (\hat{a}^\dagger)^2$   
(immediately indicates needed non-linear processes in materials, not  $E$ , but  $\hat{E}^2$ )

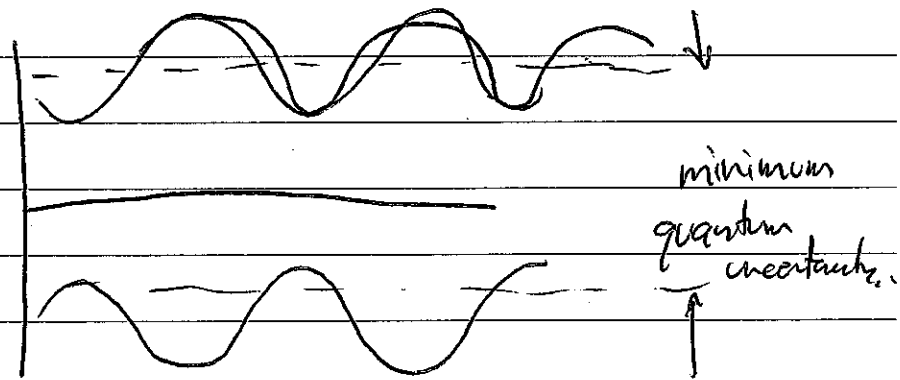
can be shown that  $g^{(2)}(\tau) = 3 + \frac{1}{\langle n \rangle}$

$\hookrightarrow$  super poisson

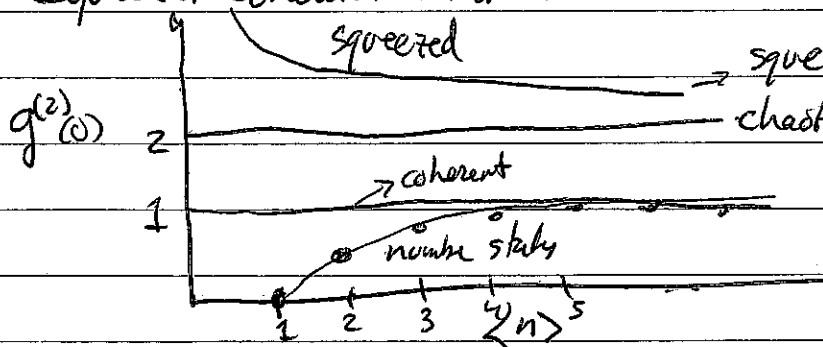


$$\langle n \rangle \neq 0 = \sinh^2 |s|$$

$\langle E \rangle$



Squeezed coherent states.



First real indication for the need for photons! (not photoelectric effect)

To deal with multiple modes (e.g. different ports on a beam splitter) or multiple frequencies (pulses of light e.g.) one needs multimode or continuous mode formalisms.

Multimode: add a  $\vec{k}$  and  $\omega_{\vec{k}}$  dependence,

$$[\hat{a}_{\vec{k}}, \hat{a}_{\vec{k}'}^\dagger] = \delta_{\vec{k}\vec{k}'}, \quad |F\rangle = |F_{k_1}\rangle |F_{k_2}\rangle |F_{k_3}\rangle \dots |F_{k_n}\rangle$$

→ get very similar results for "classical"-like light.

Continuous mode: frequency spectrum continuous:  
 imagine a box in space (size  $L$ ) with  
 frequency spacing  $\Delta\omega = \frac{2\pi c}{L}$   
 in the limit that  $L \rightarrow \infty, \Delta\omega \rightarrow 0$

example  
 for co-propagating  
 beams.

$$\sum_{\vec{k}} \rightarrow \frac{1}{\Delta\omega} \int d\omega \quad \delta_{\vec{k}\vec{k}'} \rightarrow \Delta\omega \delta(\omega - \omega')$$

$$\hat{a}_{\vec{k}} \rightarrow \sqrt{\kappa(\omega)} \hat{a}(\omega)$$

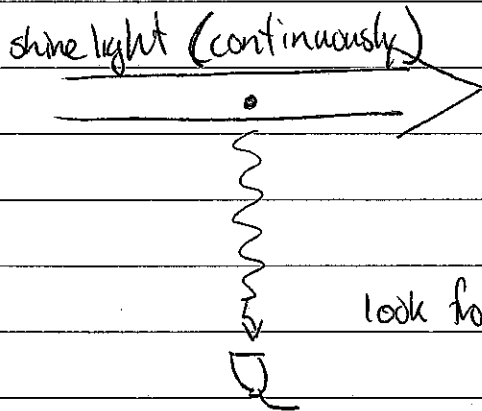
$$[\hat{a}(\omega), \hat{a}^\dagger(\omega')] = \delta(\omega - \omega')$$

eg.

$$H_R = \sum_{\vec{k}} \hbar \omega_{\vec{k}} (\hat{a}_{\vec{k}}^\dagger \hat{a}_{\vec{k}}) + \text{vacuum} \rightarrow \int_0^\infty d\omega \hbar \omega \hat{a}^\dagger(\omega) \hat{a}(\omega) + \text{vacuum.}$$

$$\hat{n} = \int d\omega \hat{a}^\dagger(\omega) \hat{a}(\omega), \quad \text{can define } \hat{a}(t) = \int_{-\infty}^{\infty} \frac{d\omega}{\sqrt{2\pi}} \hat{a}(\omega) e^{-i\omega t}$$

# Resonance Fluorescence from a single atom.



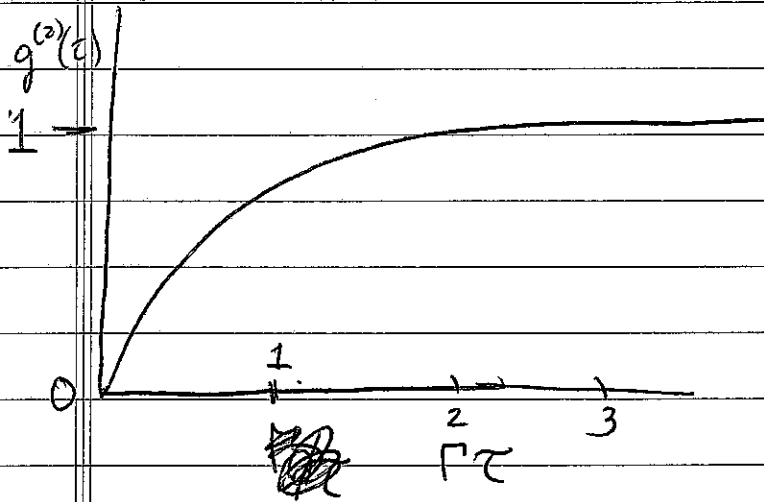
classically, the radiation is deterministic sine-wave.

quantum mechanically:

look from the side. antibunching, sub poissonian light.

$$g^{(2)}(\tau) = 1 - e^{-\Gamma/\tau}$$

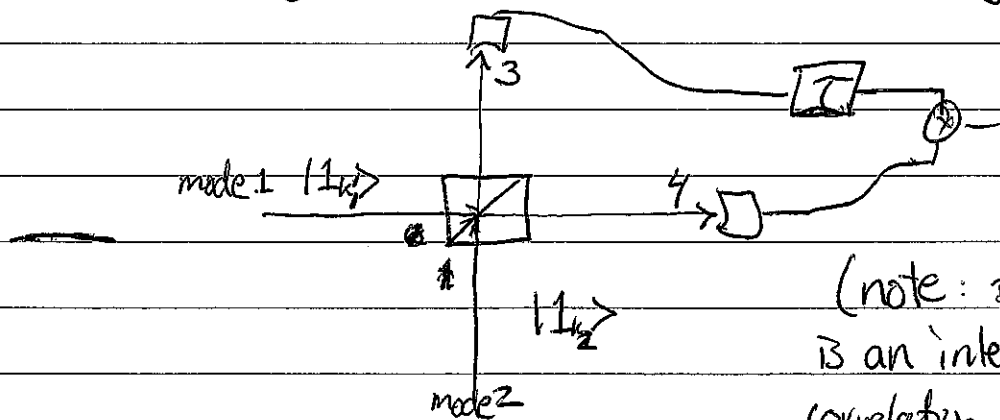
(looks like ~~collision~~ collision broadened light, but with - sign)



Takes a time  $1/\Gamma$  for the atom to get excited, decay again.

distinctly quantum effect.

# Example of 2-photon interference, Hong-Ou-Mandel



(note: photon correlation is an intensity (energy) correlation, ~~at~~  $g^{(2)}$ )

What do you get?

independently

$$|1_{k_2}\rangle \rightarrow R_{24} |1_{k_4}\rangle + T_{23} |1_{k_3}\rangle$$

$$|1_{k_1}\rangle \rightarrow R_{13} |1_{k_3}\rangle + T_{14} |1_{k_4}\rangle$$

Energy conservation gives  $|R_{24}| = |R_{13}| \equiv R$      $R^2 + T^2 = 1$   
 $|T_{13}| = |T_{14}| \equiv T$

energy conservation arguments must be true.

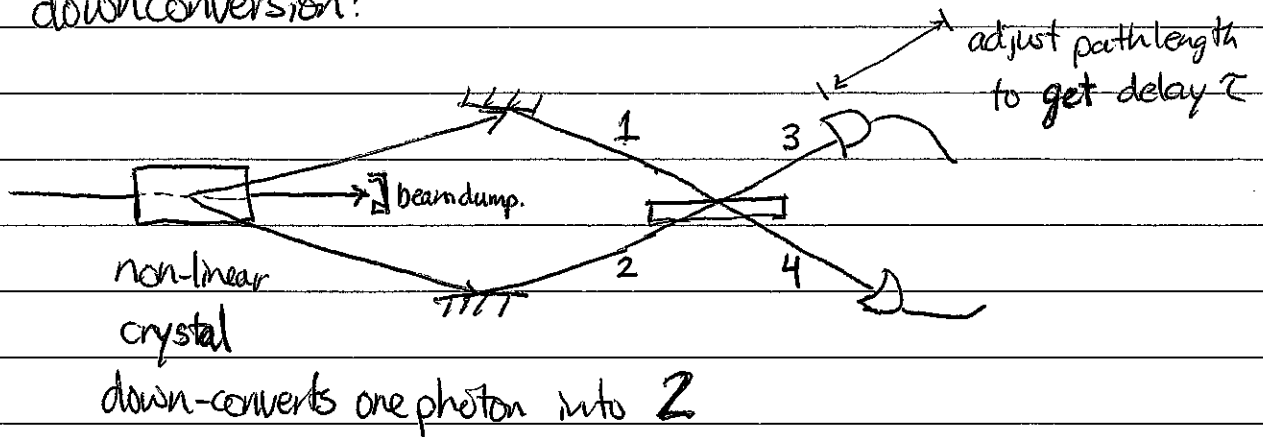
In addition, the phase shifts on reflection have to satisfy:

$$\phi_{13} + \phi_{24} - \phi_{123} - \phi_{144} = \pm \pi$$

must have the correct relative magnitude and phase to interfere and conserve energy.

example:  $R_{13} = R$      $R_{24} = -R$  } what breaks the symmetry?  
 $T_{14} = T$      $T_{23} = T$

Experimentally realized via parametric downconversion:



$$|1_{k_1}\rangle |1_{k_2}\rangle \rightarrow \cancel{|1_{k_3}\rangle} + \cancel{|1_{k_4}\rangle}$$

$$= \hat{a}_1^\dagger \hat{a}_2^\dagger |0\rangle \rightarrow \text{create photon in each input arm.}$$

The reflection/transmission properties of the beam splitter are

$$\left. \begin{aligned} \hat{a}_3^\dagger &= R\hat{a}_1^\dagger + T\hat{a}_2^\dagger \\ \hat{a}_4^\dagger &= -R\hat{a}_2^\dagger + T\hat{a}_1^\dagger \end{aligned} \right\} \Rightarrow \begin{aligned} \hat{a}_1^\dagger &= R\hat{a}_3^\dagger + T\hat{a}_4^\dagger \\ \hat{a}_2^\dagger &= T\hat{a}_3^\dagger - R\hat{a}_4^\dagger \end{aligned}$$

$$\begin{aligned} \hat{a}_1^\dagger \hat{a}_2^\dagger |0\rangle &= (R\hat{a}_3^\dagger + T\hat{a}_4^\dagger)(T\hat{a}_3^\dagger - R\hat{a}_4^\dagger) |0\rangle \\ &= RT(\hat{a}_3^\dagger)^2 + T^2\hat{a}_4^\dagger \hat{a}_3^\dagger - R^2\hat{a}_3^\dagger \hat{a}_4^\dagger + RT(\hat{a}_4^\dagger)^2 |0\rangle \\ &= \sqrt{2}RT |2_3 0_4\rangle + (T^2 - R^2) |1_3 1_4\rangle - RT\sqrt{2} |0_3 2_4\rangle \end{aligned}$$

for  $|T|^2 = |R|^2 = \frac{1}{2}$

$$|\psi_f\rangle = \frac{1}{\sqrt{2}} |2_3 0_4\rangle - \frac{1}{\sqrt{2}} |0_3 2_4\rangle$$

$$P(2_3, 0_4) = |\langle 2_3 0_4 | \psi_f \rangle|^2 = \frac{1}{2} |\langle 2_3 0_4 | 2_3 0_4 \rangle|^2 = \frac{1}{2} = P(0_3, 2_4)$$

$$P(1, 1) = 0$$

The photons (completely independent<sup>sources</sup>) always come out on the same port!

If you factor in a coherence length, ~~or~~ or if the sources are pulsed with a width  $\Delta$

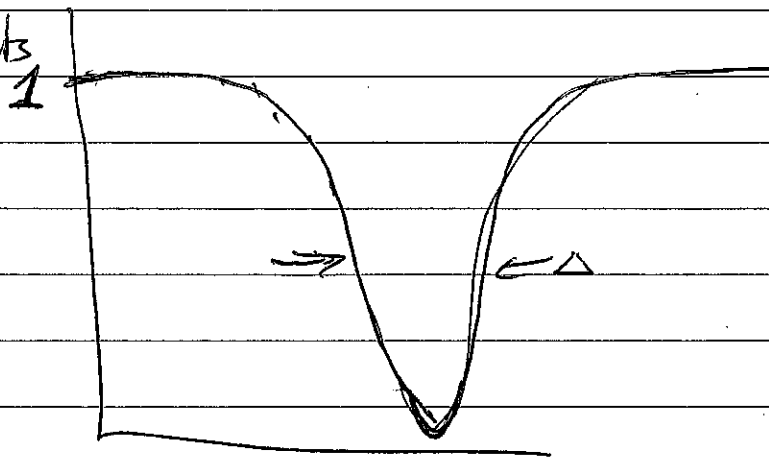
$$P(z_3, a_4) = P(o_3, z_4) = \frac{1}{4} R^2 T^2 (1 + |J|^2) \Rightarrow \frac{1}{4} (1 + |J|^2)$$

$$P(i_3, i_4) = \frac{1}{4} (1 - 2 R^2 T^2 (1 + |J|^2)) \Rightarrow 1 - \frac{1}{2} (1 + |J|^2)$$

$$J^2 = \left| \int A_1(t) A_2(t) dt \right|^2 \quad A_i = \text{envelope function.}$$

↳ overlap integral  $\cdot e^{-\frac{1}{2} \Delta^2 (t-t_c)^2}$   
for Gaussian pulses.

normalized coincidence counts



similar for a correlation width  $\rightarrow$  very non-classical.  
The classical beam splitter response is ~~is~~ Binomial

$$P(z_3, a_4) = \frac{1}{4} \quad P(i_3, i_4) = \frac{1}{4} \times 2 \quad P(o_3, z_4) = \frac{1}{4}$$