

# Matter wave Equation

- What we know:

Einstein

$$E = hf = \left(\frac{h}{2\pi}\right) (2\pi f) = \hbar \omega$$

de Broglie

$$p = \frac{h}{\lambda} = \left(\frac{h}{2\pi}\right) \left(\frac{2\pi}{\lambda}\right) = \hbar k$$

- plane wave solutions  $\Psi(x,t) \approx e^{i(kx - \omega t)}$

- Correspondence w/ classical relation  $E = p^2/2m + V$  (total = kinetic + potential)

$\Rightarrow$  Interpret  $E, p$  (and all other observables) as linear operators

Measurement yields eigenvalue of appropriate operator (more on this later)

If plane wave is eigenfunction of energy and momentum,

$$\overset{\leftrightarrow}{E} \Psi = \hbar \omega \Psi \quad \text{so} \quad \overset{\leftrightarrow}{E} = i\hbar \frac{\partial}{\partial t} \quad / \quad \overset{\leftrightarrow}{p} \Psi = \hbar k \Psi \quad \text{so} \quad \overset{\leftrightarrow}{p} = \frac{\hbar}{i} \frac{\partial}{\partial x}$$

then  $\overset{\leftrightarrow}{E} \Psi = \left(\frac{\overset{\leftrightarrow}{p}^2}{2m} + V\right) \Psi$  becomes

$$i\hbar \frac{\partial}{\partial t} \Psi(x,t) = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} \Psi(x,t) + V(x) \Psi(x,t)$$

the "time-dependent Schrodinger equation"

Simplest case:  $V(x) = 0$  "Free particle"

Solution: plane wave  $\Psi(x,t) = e^{i(kx - \omega t)}$

$$-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} \Psi = i\hbar \frac{\partial}{\partial t} \Psi$$

$$\frac{\hbar^2 k^2}{2m} \Psi = \hbar \omega \Psi \quad \Rightarrow \quad \omega = \frac{\hbar k^2}{2m} \quad \text{"dispersion relation"}$$

C.f. classical wave:  $\omega = kc$  ( $\lambda f = c$ )  $V_{\text{phase}} = \frac{\omega}{k} = c$   $V_{\text{group}} = \frac{d\omega}{dk} = c$

For matter waves,  $V_{\text{phase}} = \frac{\omega}{k} = \frac{\hbar k}{2m}$ ,  $V_{\text{group}} = \frac{d\omega}{dk} = \frac{\hbar k}{m} = \frac{p}{m} = \frac{mv}{m} = v$   
Corresponds to classical velocity!

What is  $\Psi$ ?

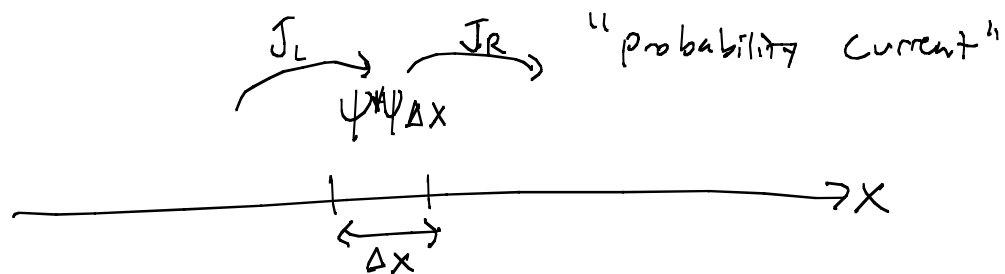
"Born Rule" / "Copenhagen Interpretation"

$|\Psi|^2 = \Psi^* \Psi$  is a probability density

Consequences:

- Normalization:  $\int_{-\infty}^{\infty} \Psi^* \Psi dx = 1$

- Probability is conserved:



$$\frac{\partial}{\partial t} (\Psi^* \Psi \Delta x) = J_L - J_R \quad \rightarrow \quad \frac{\partial}{\partial t} \Psi^* \Psi = \frac{J_L - J_R}{\Delta x}$$

In the continuous limit  $\Delta x \rightarrow 0$ ,  $\frac{\partial}{\partial t} \Psi^* \Psi = - \frac{\partial J}{\partial x}$  "continuity equation"

What is Probability current  $J$ ?

$$\frac{\partial}{\partial t} (\psi^* \psi) = \left( \frac{\partial}{\partial t} \psi^* \right) \psi + \psi^* \left( \frac{\partial}{\partial t} \psi \right)$$

From TDSE:  $\frac{\partial \psi}{\partial t} = -\frac{\hbar}{2mi} \frac{\partial^2 \psi}{\partial x^2} + \frac{V}{i\hbar} \psi$  and  $\frac{\partial \psi^*}{\partial t} = \frac{\hbar}{2mi} \frac{\partial^2 \psi^*}{\partial x^2} - \frac{V}{i\hbar} \psi^*$

$$\frac{\partial}{\partial t} (\psi^* \psi) = \frac{\hbar}{2mi} \frac{\partial^2 \psi^*}{\partial x^2} \psi - \frac{V}{i\hbar} \psi^* \psi - \psi^* \frac{\hbar}{2mi} \frac{\partial^2 \psi}{\partial x^2} + \psi^* \frac{V}{i\hbar} \psi$$

$$= -\frac{\partial}{\partial x} \left[ \frac{\hbar}{2mi} \left( \psi^* \frac{\partial \psi}{\partial x} - \frac{\partial \psi^*}{\partial x} \psi \right) \right] = -\frac{\partial J}{\partial x}$$

So  $J = \frac{\hbar}{2mi} \left( \psi^* \frac{\partial \psi}{\partial x} - \frac{\partial \psi^*}{\partial x} \psi \right) = \text{Im} \left\{ \frac{\hbar}{m} \psi^* \frac{\partial \psi}{\partial x} \right\}$

Example: free particle plane wave  $\psi = e^{i(kx - \omega t)}$

$$J = \text{Im} \left\{ \frac{\hbar}{m} e^{-i(kx - \omega t)} ik e^{i(kx - \omega t)} \right\} = \psi^* \psi \frac{\hbar k}{m} \sim \text{density} \cdot \text{velocity}$$

c.f. classical particle flux!

## Comparison with energy flux carried by classical EM wave

"Poynting" energy flux:  $\vec{j} = \frac{1}{2} \text{Re} \{ \vec{E} \times \vec{B}^* \}$

Faraday's Law:  $\vec{\nabla} \times \vec{E} = -\frac{\partial \vec{B}}{\partial t} = i\omega \vec{B}$  when time dependence of  $\vec{B}$  is  $\propto e^{i\omega t}$  as in a propagating wave. Then,  $B^* = -\frac{1}{i\omega} \vec{\nabla} \times \vec{E}^*$

If  $\vec{E} = E_z \hat{z}$  as in a wave propagating in the  $\hat{x}-\hat{y}$  plane,

$$\nabla \times E^* = \begin{vmatrix} \hat{x} & \hat{y} & \hat{z} \\ \frac{\partial}{\partial x} & \frac{\partial}{\partial y} & \frac{\partial}{\partial z} \\ 0 & 0 & E_z^* \end{vmatrix} = \hat{x} \frac{\partial E_z^*}{\partial y} - \hat{y} \frac{\partial E_z^*}{\partial x} = -(\hat{z} \times \hat{y}) \frac{\partial E_z^*}{\partial x} - (\hat{z} \times \hat{x}) \frac{\partial E_z^*}{\partial y} = -\hat{z} \times \vec{\nabla} E_z^*$$

Therefore,  $B^* = -\frac{1}{i\omega} \vec{\nabla} \times E^* = +\frac{1}{i\omega} \hat{z} \times \vec{\nabla} E_z^*$  and

$$\vec{j} = \frac{1}{2} \text{Re} \left\{ \vec{E} \times \frac{1}{i\omega} \hat{z} \times \vec{\nabla} E_z^* \right\}$$

Since  $\vec{A} \times \vec{B} \times \vec{C} = \vec{B}(\vec{A} \cdot \vec{C}) - \vec{C}(\vec{A} \cdot \vec{B})$ , this can be written

$$= -\frac{1}{2\omega} \operatorname{Re} \left\{ i \left( \hat{z} \left( \frac{\partial \epsilon_z}{\partial z} \vec{\nabla} \epsilon_z^* \right) - \vec{\nabla} \epsilon_z^* \left( \hat{z} \epsilon_z \right) \right) \right\}$$

The first term vanishes because of Gauss' Law:

$$\hat{z} \cdot \vec{\nabla} \epsilon_z^* = \frac{d}{dz} \epsilon_z^* \quad \text{but since } \vec{E} = \epsilon_z \hat{z}, \quad \vec{\nabla} \cdot \vec{E}^* = \frac{\partial}{\partial z} \epsilon_z^* = 0$$

Then

$$\vec{J} = \frac{1}{2\omega} \operatorname{Re} \left\{ i \epsilon_z \vec{\nabla} \epsilon_z^* \right\} = \frac{1}{2\omega} \operatorname{Im} \left\{ \epsilon_z^* \vec{\nabla} \epsilon_z \right\} \quad (\text{energy flux in } \mathcal{E} + \mathcal{M})$$

This is mathematically identical to

$$\vec{J} = \frac{\hbar}{m} \operatorname{Im} \left\{ \psi^* \vec{\nabla} \psi \right\} \quad (\text{probability flux in QM})$$