

Differential Equations

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Two important rules of D.E.s.

• The best way to solve a D.E. is to guess a solution and try it.

• A D.E. which has a maximum of (n) order derivatives has

- ① (n) free parameters AND
- ② requires (n) initial conditions for a unique solution

A few D.E. are so simple and common that you are required to know their solutions:

①  $\frac{dy(t)}{dt} = A \Rightarrow \text{soln: } y(t) = At + C$

↑  
1<sup>st</sup> order

↑  
one free parameter

Need one initial condition to specify C.

②  $\frac{dy(t)}{dt} = Ay(t) \Rightarrow \text{soln: } y(t) = Ce^{At}$

↑  
1<sup>st</sup> order.

↑  
one free parameter

equivalent

②

$$\textcircled{5} \quad \frac{d^2 y(t)}{dt^2} + Ay(t) = 0 \Rightarrow \text{Many}^{\wedge} \text{ ways to write the solution:}$$

↳ 2nd order

Solution :  $y(t) = C_1(\sin(\sqrt{A}t) + C_2)$

OR  $y(t) = C_1(\cos(\sqrt{A}t) + C_2)$

OR  $y(t) = C_1(\sin(\sqrt{A}t)) + C_2 \cos(\sin(\sqrt{A}t))$

OR  $y(t) = C_1 e^{i\sqrt{A}t} + C_2 e^{-i\sqrt{A}t}$

OR  $\vdots$

~~Example :  $\frac{d^2 y(t)}{dt^2} + 4y(t) = 0$~~

~~Solution :~~

### Wave Formula Sheet

$\lambda$  = wavelength (meters)

$k$  = "wave number" (meters<sup>-1</sup>) =  $\frac{2\pi}{\lambda}$

$T$  = period (seconds)

$\nu$  = frequency =  $\frac{1}{T}$  = (seconds<sup>-1</sup> or Hz)

$\omega$  = angular frequency =  $\frac{2\pi}{T} = 2\pi\nu$  (Hz)

Phase difference between two points in space:

$$\Delta\phi = kx, \quad x = \text{distance between the points}$$

Phase difference after waiting for time (t):

$$\Delta\phi = 2\pi \left(\frac{t}{T}\right) = \omega t$$

Planck Formula: Light is composed of photons.

Energy of each photon is  $E = h\nu = \left(\frac{h}{2\pi}\right) (2\pi\nu) = \hbar\omega$

↑  
Planck's  
Constant

↓  
 $\hbar = \frac{h}{2\pi}$

∴  $E = \hbar\omega$  Energy of a photon

DeBroglie: Both matter particles and light particles have a wavelength:

DeBroglie Wavelengths =  $\lambda = \frac{h}{p}$

↑  
particle  
momentum

For photons, momentum =  $p = \frac{E}{c}$  from special relativity

↑  
speed of light

So  $\frac{h}{p} = \frac{h}{(E/c)} = \left(\frac{h}{E}\right) c = \left(\frac{1}{\nu}\right) c = \lambda$

So DeBroglie's wavelength makes sense for photons

If  $\lambda = \frac{h}{p}$  then  $p = \frac{h}{\lambda} = \left(\frac{h}{2\pi}\right) \left(\frac{2\pi}{\lambda}\right) = \hbar k$

Summary:  $E = h\nu = \hbar\omega$   
 $p = \hbar k$



# Postulate I of QM: Eigenvalues and Eigenfunctions

For any observable  $A$ , there exists

- ① an "operator"  $\Rightarrow \hat{A}$  ← hat means operator.
- ② "eigenvalues"  $\Rightarrow a$  ← a numerical value
- ③ "eigenfunctions"  $\Rightarrow \psi$  ← a function

These are related by:

$$\boxed{\hat{A} \psi = a \psi} \quad \text{"An eigenvalue equation"}$$

If we measure  $A$ , the result will always be an eigenvalue of  $A$ .

What is an operator? It's an instruction to do something.

<p><u>Operator</u></p> <p><math>\hat{D} \equiv \frac{d}{dx}</math></p> <p>Identity = <math>\hat{I} \equiv 1</math></p> <p><math>\hat{B} \psi</math></p>	<p><u>Instructions</u></p> <p>take the derivative : <math>\hat{D} \psi = \frac{d\psi}{dx}</math>.</p> <p>multiply by one : <math>\hat{I} \psi = \psi</math>.</p> <p>Annihilate <math>\psi</math> : <math>\hat{B} \psi = \emptyset</math>.</p>
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The two most important operators in QM are

① Momentum:  $\hat{p} = -i\hbar \frac{d}{dx}$  in 1D

② position:  $\hat{x} = \text{multiply by } x$  in 1D.

Postulate II: The state function  $\psi(x)$ .

All properties of a QM system are described by the "state function" or "wave function":  $\psi(x)$ .

Corollary ①: The expectation value for an observable  $C$  can be calculated by

$$\langle C \rangle = \int_{-\infty}^{\infty} \psi^*(x) \hat{C} \psi(x) dx$$

where  $\hat{C}$  is the operator for  $C$ .

Ex:  ~~$\langle p \rangle$~~  =  $\int_{-\infty}^{\infty} \psi^*(x) \hat{p} \psi(x) dx$

$$\langle p \rangle = -i\hbar \int_{-\infty}^{\infty} \psi^*(x) \frac{d}{dx} \psi(x) dx$$

Corollary ②:  $|\psi(x)|^2 dx = \psi^*(x) \psi(x) dx = \text{Probability}$

to find the particle between  $x$  and  $x+dx$ .  
 "Born Interpretation of the wavefunction."

With the Born Interpretation,

$$\langle x \rangle = \int_{-\infty}^{\infty} \psi^*(x) \hat{x} \psi(x) dx = \int_{-\infty}^{\infty} \psi^* x \psi dx$$

↑  
multiply  
by x

$$= \int_{-\infty}^{\infty} x |\psi|^2 dx$$

↓ Born Interpretation

This is what  
we expect

$$\langle x \rangle = \int_{-\infty}^{\infty} x P(x) dx$$

from the statistical definition of the expectation  
value of x.

Phys 401 - Lecture 6 - Fall 2011

Postulate I: For every observable  $A$ , we have

$$\boxed{\hat{A} \psi = a \psi} \leftarrow \text{eigenvalue equation}$$

↑ an operator
↑ an eigenvalue
↑ an eigenfunction

If we measure  $A$ , we always get eigenvalues of  $A$ .

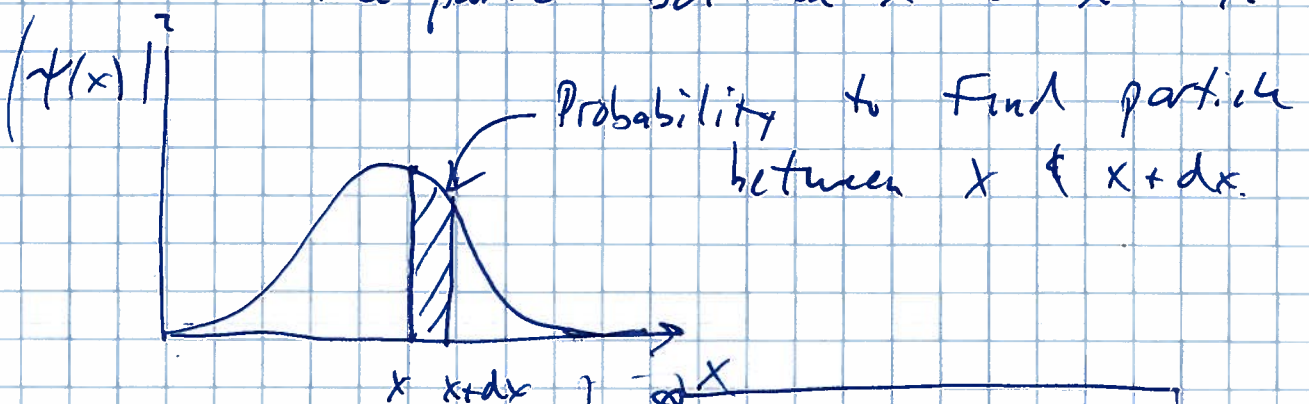
Postulate II: Statistics & Wavefunctions

$$\langle C \rangle = \int_{-\infty}^{\infty} \psi^*(x) \hat{C} \psi(x) dx$$

$$\psi^*(x) \psi(x) dx = |\psi(x)|^2 dx = P(x) dx$$

Probability to find  $\uparrow$

The particle between  $x$  and  $x+dx$ .



Also, note that

$$\int_{-\infty}^{\infty} |\psi(x)|^2 dx = 1$$

"Normalization Condition"

total probability to be anywhere

### Postulate III : - Collapse of the Wavefunction

If we measure  $A$ , and eigenvalue  $(a)$  is the result, then after the measurement the state function for the system will be  $\psi_a$ , the eigenfunction corresponding to  $(a)$ .

### Postulate IV : Time Evolution of the state function

If we make no measurements, then the state function evolves in time according to the Time-Dependent Schrödinger Eq. (TDSE)

$$i\hbar \frac{\partial}{\partial t} \Psi(x,t) = \hat{H} \Psi(x,t)$$

where  $\hat{H}$  is the energy operator for the system (the Hamiltonian.)

• Convention : we'll use capital psi  $\Psi$  for state functions which depend on time :  $\Psi(x,t)$  and lower case ~~psi~~ psi  $\psi$  for state functions which do not :  $\psi(x)$ .



Summary

$$\hat{A} \psi = a \psi$$

$$\langle A \rangle = \int_{-\infty}^{\infty} \psi^*(x) \hat{A} \psi(x) dx$$

$$|\psi(x)|^2 dx = P(x) dx$$

$$\int_{-\infty}^{\infty} |\psi(x)|^2 dx = 1$$

$$i\hbar \frac{\partial}{\partial t} \Psi(x,t) = \hat{H} \Psi(x,t)$$

Some important eigenvalue equations

Momentum :  $\hat{p} = -i\hbar \frac{d}{dx}$

Eigenvalue Eq :  $\hat{p} \psi = p_x \psi$

$$-i\hbar \frac{d}{dx} \psi = p_x \psi \leftarrow \text{A differential eq.}$$

Solution :  $\psi(x) = A e^{i(p_x/\hbar)x}$  for  $\psi(x)$ .

$$\frac{p_x}{\hbar} = k$$

$$\psi(x) = A e^{ikx}$$

The eigenfunctions for of momentum are exponential functions.

We call  $A e^{ikx}$  a plane wave. The wavelength

$$\lambda = \frac{2\pi}{k} = \frac{2\pi}{(p_x/\hbar)} = \frac{h}{p_x} \leftarrow \text{DeBroglie}$$

Ex: We measure the momentum of a particle, and the result is value  $p = \hbar k'$ .

After the measurement, the state function is

$$\psi(x) = \alpha(x) = A e^{ik'x}.$$

Energy: The Hamiltonian Operator  $\hat{H}$ .

In classical mechanics, Energy =  $\frac{1}{2}mv^2 + V(x)$

for a single particle in a potential  $V(x)$ .

We can write this as

$$\begin{aligned} E &= \frac{1}{2m} (mv)^2 + V(x) \\ &= \frac{p^2}{2m} + V(x) \end{aligned}$$

In QM, we assume that

$$\begin{aligned} \hat{H} = \text{energy operator} &= \frac{1}{2m} (\hat{p})^2 + V(x) \\ &= \frac{1}{2m} \hat{p} \hat{p} + V(x) \\ &= \frac{1}{2m} \left( -i\hbar \frac{d}{dx} \right) \left( -i\hbar \frac{d}{dx} \right) + V(x) \end{aligned}$$

$$\boxed{\hat{H} = \frac{-\hbar^2}{2m} \frac{d^2}{dx^2} + V(x)}$$

Eigenvalue Equation for  $\hat{H}$ :

$$\boxed{\hat{H} \alpha = E \alpha} \leftarrow \text{Time-Independent Schrodinger Equation}$$

(5)

Free Particle Case:  $V(x) = 0$ .Then  $\hat{H}\psi = E\psi$  i.e.

$$\frac{\hbar^2}{2m} \frac{d^2}{dx^2} \psi(x) = E \psi(x)$$

$$\psi''(x) + \frac{2mE}{\hbar^2} \psi(x) = 0$$

← A D.E. for the eigenfunctions of energy.

Solution:  $\psi(x) = A e^{ikx} + B e^{-ikx}$ 

where  $k = \frac{\sqrt{2mE}}{\hbar}$

or  $E = \frac{(\hbar k)^2}{2m} = \frac{(p_x)^2}{2m}$  as expected for the kinetic energy of a free particle.

Ex: Is  $\psi(x) = A \cos(kx)$  an eigenfunction of

① momentum?

Answer: Try it:  $\hat{p} A \cos(kx) = -i\hbar A \frac{d}{dx} \cos(kx)$

$$= +i\hbar k A \sin(kx)$$

$$\stackrel{?}{=} p_x A \cos(kx) ??$$

No -  $\sin(kx) \neq \cos(kx)$

So  $\cos(kx)$  is not an e.f. of  $\hat{p}$ .

Is  $\psi(x) = A \cos(kx)$  an eigenfunction of

the free particle Hamiltonian:  $\hat{H} = -\frac{\hbar^2}{2m} \frac{d^2}{dx^2}$  ?

Try it:  $-\frac{\hbar^2}{2m} \frac{d^2}{dx^2} A \cos(kx)$

$$= \frac{\hbar^2 k^2}{2m} A \cos(kx) \stackrel{?}{=} E A \cos(kx)$$

Yes! and the eigenvalue is

$$E = \frac{\hbar^2 k^2}{2m}$$

According to Postulate IV, the "equation of motion" of QM is the time-dependent Schrödinger Eq:

$$i\hbar \frac{\partial}{\partial t} \Psi(x,t) = \hat{H} \Psi(x,t).$$

Given a wavefunction at  $t=0$ , the TDSE tells us how it evolves in time.

Good news: We can solve this equation once and be done with it.

We guess a "separable" solution:

$$\Psi(x,t) = Q(x) T(t)$$

$Q(x)$  only depends on position.  $T(t)$  only depends on time.

Plug in our guess:

$$i\hbar Q(x) T'(t) = \hat{H} [Q(x) T(t)]$$

To simplify, assume that  $\hat{H}$  does not depend on time.

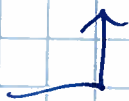
For example, assume that  $\hat{H} = \frac{-\hbar^2}{2m} \frac{d^2}{dx^2} + V(x)$ .

Then  $\hat{H} [Q(x) T(t)] = T(t) \hat{H} Q(x)$

$$\therefore i\hbar Q(x) T'(t) = T(t) \hat{H} Q(x) \quad (2)$$

$$\frac{i\hbar T'(t)}{T(t)} = \frac{\hat{H} Q(x)}{Q(x)}$$

only depends on time



only depends on x

If these are equal to each other for all  $t$  &  $x$ , they must both be equal to a constant

$$\frac{i\hbar T'(t)}{T(t)} = E = \frac{\hat{H} Q(x)}{Q(x)}$$



Separation constant, a number.

Two equations:

$$i\hbar T'(t) = E T(t) \Rightarrow \text{Solution: } T(t) = A e^{-iEt/\hbar}$$

$$\hat{H} Q(x) = E Q(x) \leftarrow \text{Time-Independent Schrödinger Eq}$$

$\therefore$  Our solution is

$$\Psi(x,t) = A Q(x) e^{-iEt/\hbar}$$

where  $E$  = energy of the system

$Q(x)$  = an energy eigenfunction of the system.

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In general, there will be many energy eigenfunctions for a system, each with its own eigenvalue.

$$\psi_1(x), \psi_2(x), \psi_3(x), \dots$$

$$E_1, E_2, E_3, \dots$$

Each of these is a solution to the TDSE:

$$\Psi_1(x,t) = A_1 \psi_1(x) e^{-i(E_1/\hbar)t}$$

$$\Psi_2(x,t) = A_2 \psi_2(x) e^{-i(E_2/\hbar)t}$$

$$\Psi_3(x,t) = A_3 \psi_3(x) e^{-i(E_3/\hbar)t}$$

$$\vdots$$

Question: What if we add together some of these solutions? Is the sum a solution to the TDSE?

Answer: Try it!

$$\text{Let } \Psi(x,t) = \sum_{n=1}^{\infty} a_n \psi_n(x) e^{-iE_n t/\hbar}$$

↑  
any set of coefficients  
at all!

Plug it in:

$$i\hbar \frac{\partial}{\partial t} \Psi(x,t) \stackrel{?}{=} \hat{H} \Psi(x,t)$$

$$i\hbar \frac{\partial}{\partial t} \left[ \sum_{n=1}^{\infty} a_n \psi_n(x) e^{-iE_n t/\hbar} \right] \stackrel{?}{=} \hat{H} \left[ \sum_{n=1}^{\infty} a_n \psi_n(x) e^{-iE_n t/\hbar} \right]$$

$$i\hbar \sum_{n=1}^{\infty} a_n \psi_n(x) \frac{\partial}{\partial t} \left[ e^{-iE_n t/\hbar} \right] \stackrel{?}{=} \sum_{n=1}^{\infty} a_n e^{-iE_n t/\hbar} \hat{H} \left[ \psi_n(x) \right]$$

$\frac{-iE_n}{\hbar} e^{-iE_n t/\hbar}$ 
 $E_n \psi_n(x)$

because  $\psi_n$  is an e.f. of  $\hat{H}$

$$\sum_{n=1}^{\infty} a_n e^{-iE_n t/\hbar} E_n \psi_n(x) \stackrel{\checkmark}{=} \sum_{n=1}^{\infty} a_n e^{-iE_n t/\hbar} E_n \psi_n(x)$$

Yes!

Any linear combination of separable solutions is a solution

This is the principle of superposition. In QM, we can add two ~~solutions~~ valid states together and the result is another valid state.



Our strategy with every system in QM will be:

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① ~~Find~~ Guess the Hamiltonian for our system.

(Usually  $\hat{H} = \frac{-\hbar^2}{2m} \frac{d^2}{dx^2} + V(x)$  in Phys 401)

② Find the energy eigenfunctions for our  $\hat{H}$  by solving the TIDE:

$$\hat{H} \psi(x) = E \psi(x)$$

Solutions:  $\psi_1(x), \psi_2(x), \psi_3(x), \psi_4(x) \dots$

$E_1, E_2, E_3, E_4, \dots$

③ ~~For~~ If we know that at  $t=0$ , the wavefunction is  $\Psi(x, t=0) = \psi(x)$ , then to get  $\Psi(x, t)$  for all time do this:

Find the set of coefficients  $\{a_n\}$  such that

$$\psi(x) = \sum_{n=1}^{\infty} a_n \psi_n(x).$$

$$\text{Then } \Psi(x, t) = \sum_{n=1}^{\infty} a_n \psi_n(x) e^{-iE_n t/\hbar}$$