

Hydrogen Atom

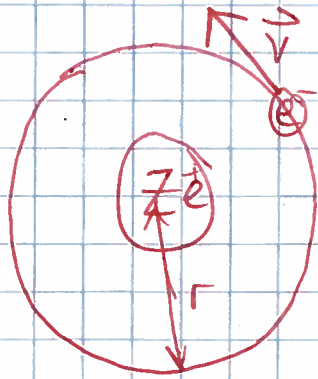
Rydberg Formula (empirical)

Atomic hydrogen emits light with a discrete set of wavelengths.

$$\frac{1}{\lambda} = R \left(\frac{1}{n_1^2} - \frac{1}{n_2^2} \right) \quad \text{where } n_1 \text{ \& } n_2 \text{ are integers.}$$

$$R = \text{Rydberg constant} \sim 1.097 \times 10^7 \text{ m}^{-1}$$

Bohr: A ~~planetary~~ planetary model of atom:



$$\text{Classical Picture: } F = \frac{Zke^2}{r^2}$$

$$\text{Acceleration} = \frac{v^2}{r} \quad (\text{centripetal})$$

$$F = ma \Rightarrow \boxed{\frac{Zke^2}{r^2} = \frac{mv^2}{r}} \Rightarrow \boxed{\frac{Zke^2}{r} = mv^2}$$

$$\text{Also } \frac{Zke^2}{2r} = \frac{1}{2}mv^2 = KE$$

$$U = -\frac{ke^2 Z}{r}$$

$$\therefore \text{Total energy} = KE + U = -\frac{ke^2 Z}{2r}$$

Angular Momentum: $L = mvr$

(2)

Bohr Assumed that L is quantized.

$$L \stackrel{?}{=} n\hbar, \quad n = \text{integer (Bohr guess)}$$

Actually, $L = \sqrt{\hbar^2 l(l+1)} = \hbar \sqrt{l(l+1)}, \quad l > \text{integer},$

So Bohr guessed incorrectly.

Nevertheless, with Bohr assumption,

$$L = mvr = n\hbar \quad \Rightarrow \quad r_n = \frac{n\hbar}{mv} \quad v = \frac{n\hbar}{mr}$$

~~$$E = KE + U = \frac{1}{2}mv^2 - \frac{ke^2Z}{r} = \frac{1}{2} \frac{n^2\hbar^2}{m^2r^2} - \frac{ke^2Z}{r}$$~~

$$\therefore \frac{Zke^2}{r} = \frac{m^2\hbar^2}{nr^2}$$

$$r = \frac{n^2\hbar^2}{mZke^2} \equiv \frac{n^2}{Z} a_0$$

where $a_0 = \text{Bohr radius}$
 $= \frac{\hbar^2}{mke^2}$

~~$$KE = \frac{1}{2}mv^2 = \frac{1}{2} \frac{n^2\hbar^2}{mr^2}$$~~
$$\text{Total Energy} = KE + U = -\frac{ke^2Z}{2r} = -\frac{ke^2Z}{2(n^2a_0)}$$

$$= -\frac{ke^2Z^2}{2a_0} \frac{1}{n^2} \rightarrow 13.6 \text{ eV for } Z=1$$

If electrons transition from n_2 to n_1 , then they emit light with energy

$$\Delta E = \frac{-ke^2Z^2}{2a} \left(\frac{1}{n_2^2} - \frac{1}{n_1^2} \right) = h\nu = \frac{hc}{\lambda}$$

$$\frac{1}{\lambda} = \frac{13.6 \text{ eV}}{hc} \left(\frac{1}{n_2^2} - \frac{1}{n_1^2} \right)$$

$$= \frac{13.6 \text{ eV}}{(1240 \text{ eV}\cdot\text{nm})} \left(\frac{1}{n_2^2} - \frac{1}{n_1^2} \right)$$

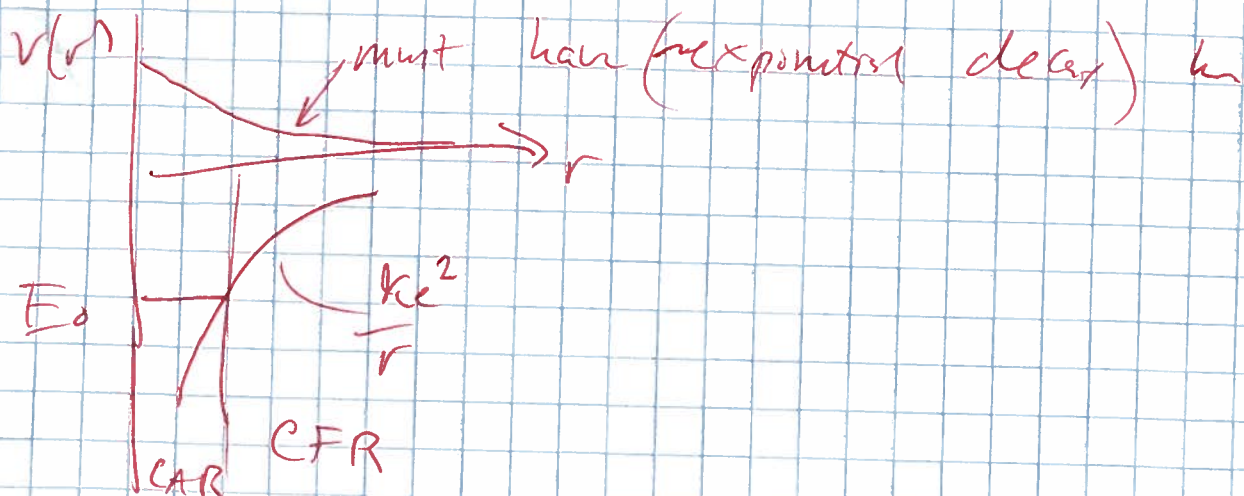
$$= \frac{13.6 \text{ eV}}{(1.24 \times 10^{-4} \text{ eV}\cdot\text{cm})} \left(\frac{1}{n_2^2} - \frac{1}{n_1^2} \right)$$

$$= 1.097 \text{ cm}^{-1} \left(\frac{1}{n_2^2} - \frac{1}{n_1^2} \right)$$

Bohr's semiclassical approach gives the correct result for the energy levels of H atoms, He⁺ ion, etc, even though his angular momentum guess was wrong.

(4)

With the full QM, energy eigenstates are quantized for bound states ~~because~~ due to the normalization condition on $|\psi|^2$.



Since $V(r)$ is a central potential, we know that

$$\phi(r, \theta, \varphi) = R(r) Y_l^m(\theta, \varphi)$$

and $R(r) = \text{radial wavefunction} = \frac{u(r)}{r}$

when $u(r)$ satisfies the radial equation:

$$-\frac{\hbar^2}{2m} \frac{d^2 u}{dr^2} + \left[V(r) + \frac{\hbar^2 l(l+1)}{2mr^2} \right] u(r) = E u(r)$$

with $V(r) = -\frac{kZe^2}{r}$

Simplify notation: $\kappa \equiv \frac{\sqrt{-2mE}}{\hbar}$ (E is negative)

$\rho \equiv \kappa r$ (dimensionless) distance

$\rho_0 \equiv \frac{2mZe^2k}{\hbar^2 \kappa}$ [note that ρ_0 determines E : $E =$

The $\frac{d^2 u}{d\rho^2} = \left[1 - \frac{\rho_0}{\rho} + \frac{l(l+1)}{\rho^2} \right] u$

Guess solutions

$u(\rho) = \rho^{l+1} e^{-\rho} \left[\sum_{j=0}^{\infty} c_j \rho^j \right]$

Some power series polynomial. The question is: what are the c_j ?

Answer: substitute the guess into the radial eq:

Result:

$$\sum_{j=0}^{\infty} j(j+1) c_{j+1} \rho^j + 2(l+1) \sum_{j=0}^{\infty} (j+1) c_{j+1} \rho^j - 2 \sum_{j=0}^{\infty} j c_j \rho^j + [\rho_0 - 2(l+1)] \sum_{j=0}^{\infty} c_j \rho^j = 0$$

For this to equal zero, each power of ρ^j must have zero as its coefficient. This means

$$j(j+1)c_{j+1} + 2(\lambda+1)(j+1)c_{j+1} - 2jc_j + [p_0 - 2(\lambda+1)]c_j = 0$$

$$\text{or } c_{j+1} = \frac{2(j+\lambda+1) - p_0}{(j+1)(j+2\lambda+2)} c_j \leftarrow \begin{array}{l} \text{Recursion} \\ \text{Formula} \\ \text{for } c_j \end{array}$$

Given c_0 , this formula gives c_1 .

Then c_1 gives c_2 .

Then c_2 gives c_3 .

etc.

But this is dangerous, because then the series might go to infinity as $\rho \rightarrow \infty$ (This means $r \rightarrow \infty$).

There ~~can~~ ^{would be} a normalizable solution

To make sure the radial wavefunction is normalizable, we must have some coefficient $c_{j_{\max}}$

where $c_{j_{\max}+1} = 0$, so that

$$c_{j_{\max}+1} = 0$$

$$c_{j_{\max}+2} = 0$$

Then $u(r)$ will be

normalizable