

Schedule

Today: Begin Lab 6

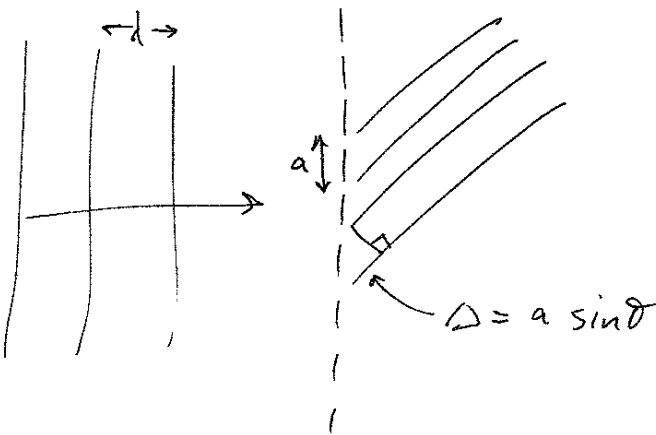
11/27: Finish Lab 6 - turn in homework # 7

12/4: Exam Review - Lab 6 due

12/11: Final Exam?

12/17: Final Exam

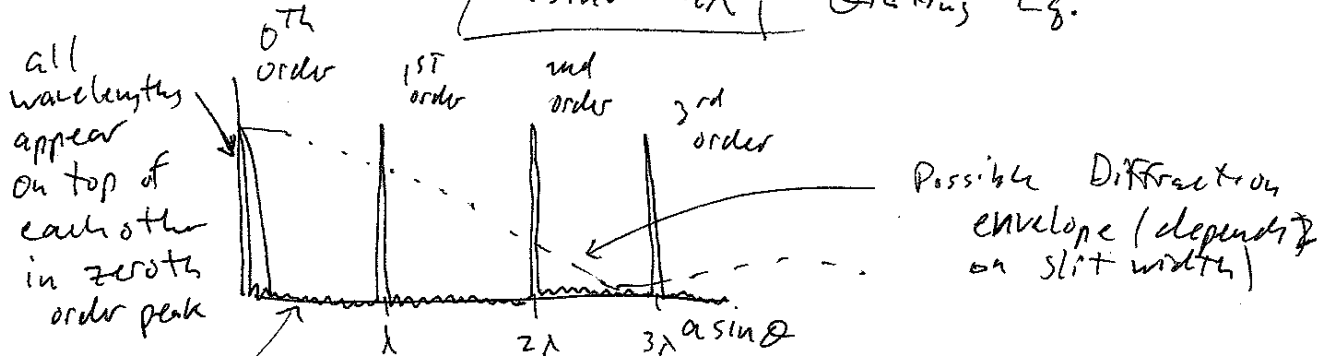
Diffraction Grating



Maxima occur when $\Delta = n\lambda$

$a \sin \theta = n\lambda$

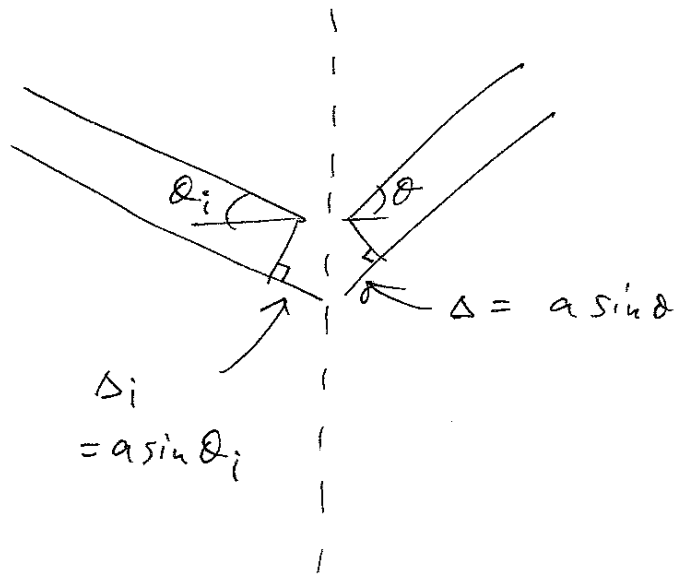
 Grating Eq.



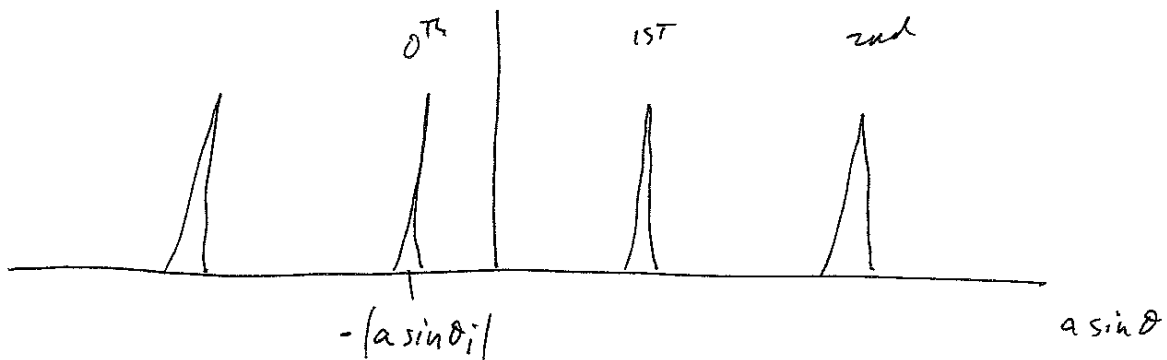
Today's lab:

- use known wavelength of Hg lamp to measure grating spacing a
- use measured grating spacing to measure wavelength of the lamp & Al_K source.

Non-zero angle of incidence



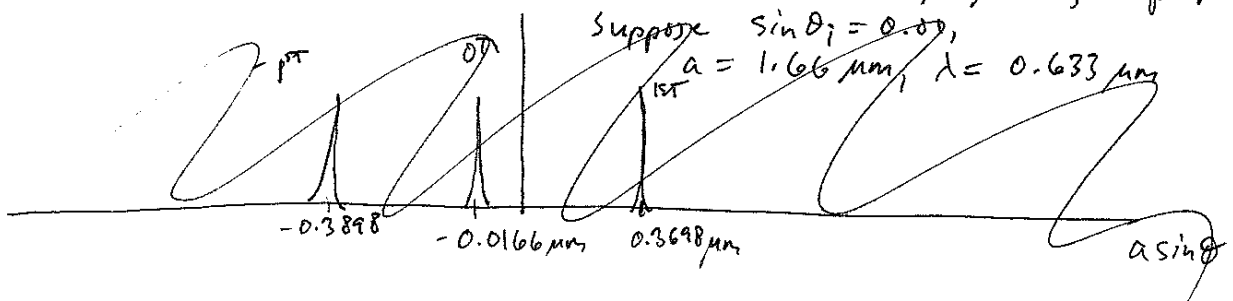
maxima when $a (\sin \theta_i + \sin \theta) = n \lambda$



θ is measured from the normal to the grating.

But we only know roughly what angle corresponds to normal.

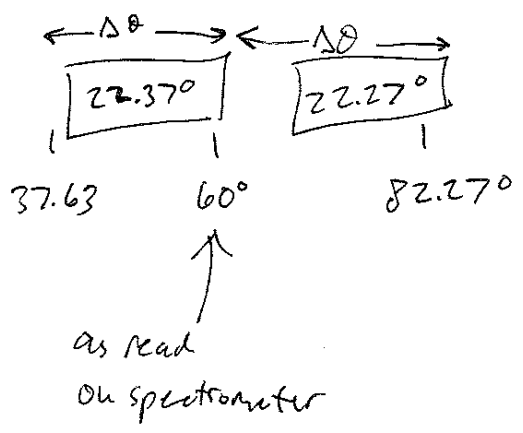
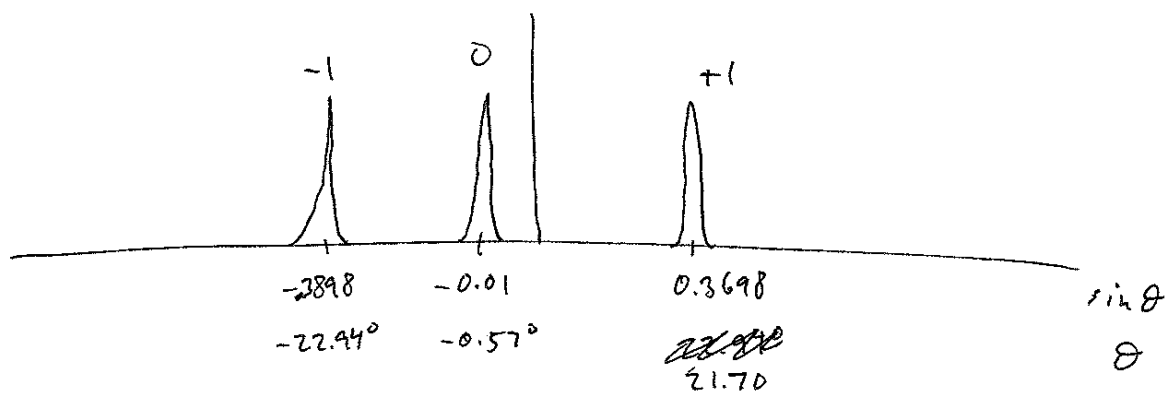
\Rightarrow Measure angular position of 1^{st} & -1^{st} order peaks. When $a \sin \theta (n=1) = -a \sin \theta (n=-1)$, grating is perpendicular.



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Suppose $a = 1.66 \mu\text{m}$, $\lambda = 0.633 \mu\text{m}$, and $\theta_i = 0.01$



Differ by ~~0.1~~ 0.1° due to $\theta_i \neq 0$.

\Rightarrow Adjust angle of grating until +1 and -1 peaks appear at equal angular distance from $\theta = \theta_i$.

Hydrogen Energy Levels

3

$$E_n = \frac{-13.6 \text{ eV}}{n^2}, \quad n = 1, 2, 3, \dots$$

$$\Delta E = -13.6 \text{ eV} \left(\frac{1}{n^2} - \frac{1}{n'^2} \right) = h\nu$$

frequency.

↑
Planck's
constant

$$\frac{1}{\lambda} \equiv \tilde{\nu} = \text{wavenumber}$$
$$= \frac{\nu}{c} = \frac{\Delta E}{hc} = R \left(\frac{1}{n^2} - \frac{1}{n'^2} \right)$$

$$hc = 1240 \text{ eV} \cdot \text{nm}$$

R = Rydberg Constant, units of cm^{-1}

For transitions to $n=1$, ν is UV (invisible) (Lyman Series)

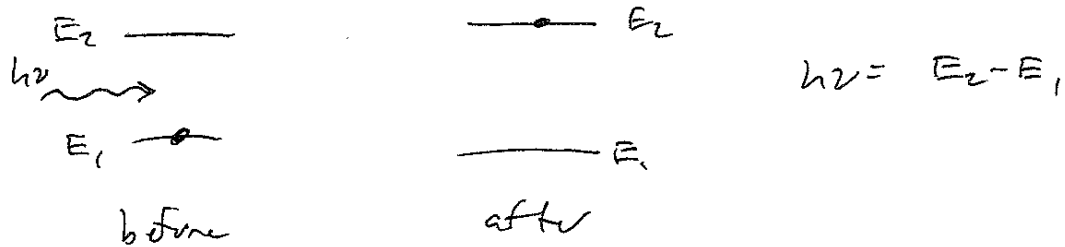
Transition to $n=2$, ν is visible (Balmer Series)

- Measure Balmer series wavelengths
- Guess which wavelength correspond to which n'
- Extract Rydberg Constant.

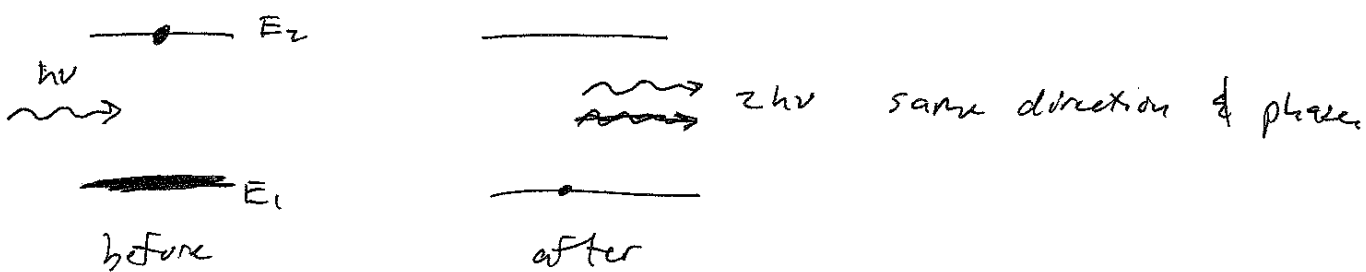
Lasers

Three interactions between light and atoms:

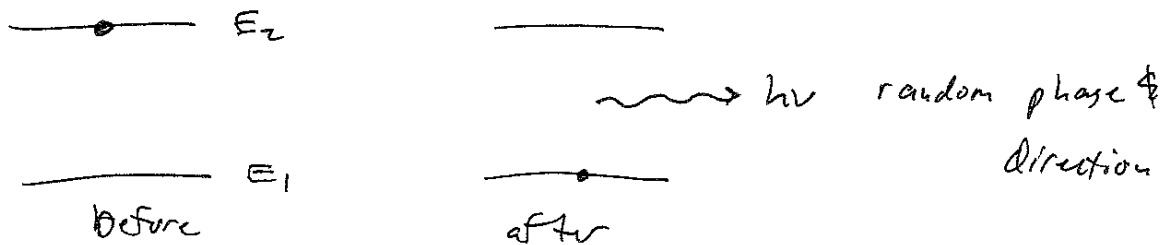
• Stimulated absorption



• Stimulated emission



• Spontaneous emission

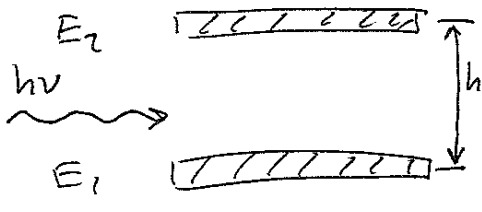


Conventional light sources rely on spontaneous emission.

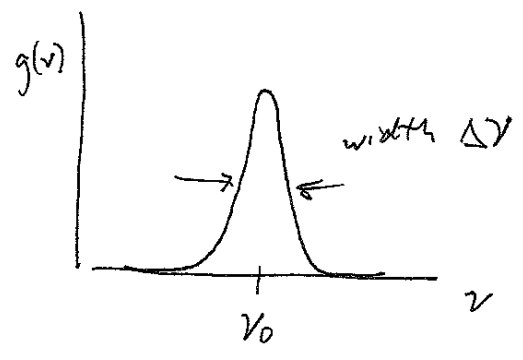
Lasers use stimulated emission.

Lineshape Function

Real atomic energy levels have a small spread:



The probability that a photon will interact with the atom is called the lineshape function, $g(\nu)$



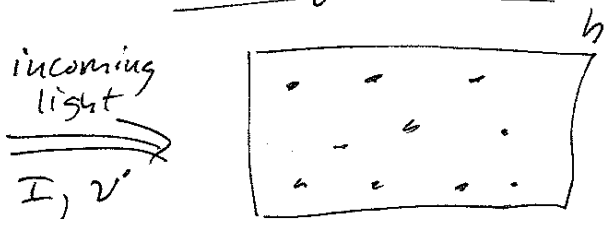
Typically for an atom in vacuum,

$$\frac{\Delta\nu}{\nu_0} \approx 10^{-8} \text{ to } 10^{-5}$$

$$\int g(\nu) d\nu = 1 \text{ (normalized)}$$

→ Stop Here

Rate Equations



N_1 in energy level E_1
 N_2 in energy level E_2

Rate of Stimulated Absorption per unit volume:

$$R_{\text{stim. abs}} = B_{12} g(\nu') \left(\frac{I}{c}\right) N_1$$

↑ proportionality constant (Einstein coeff.)

Rate of Stimulated Emission per unit volume:

$$R_{\text{stim. em}} = B_{21} g(\nu') \left(\frac{I}{c}\right) N_2$$

↑ Einstein coeff.

Rate of Spontaneous Emission per unit volume:

$$R_{\text{spont. emiss}} = A_{21} N_2 \text{ Einstein coeff.}$$

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A_{21} , B_{12} , B_{21} can be calculated with quantum mechanics.

Einstein showed that in thermal equilibrium,

$$\frac{A_{21}}{B_{21}} = \frac{8\pi h\nu^3}{c^3} \quad \text{and} \quad B_{21} = B_{12}$$

But these ratios must also be true out of thermal equilibrium, because A_{21} , B_{12} , B_{21} are fundamental properties of the atoms.