

# HW 6 Solutions

$$34.3 \quad \frac{E}{B} = c \quad \text{or} \quad \frac{220}{B} = 3.00 \times 10^8; \quad \text{so} \quad B = 7.33 \times 10^{-7} \text{ T} = \boxed{733 \text{ nT}}$$

$$34.4 \quad \frac{E_{\max}}{B_{\max}} = v \text{ is the generalized version of Equation 34.13.}$$

$$B_{\max} = \frac{E_{\max}}{v} = \frac{7.60 \times 10^{-3} \text{ V/m}}{(2/3)(3.00 \times 10^8 \text{ m/s})} \left( \frac{\text{N} \cdot \text{m}}{\text{V} \cdot \text{C}} \right) \left( \frac{\text{T} \cdot \text{C} \cdot \text{m}}{\text{N} \cdot \text{s}} \right) = 3.80 \times 10^{-11} \text{ T} = \boxed{38.0 \text{ pT}}$$

$$34.7 \quad (a) \quad B = \frac{E}{c} = \frac{100 \text{ V/m}}{3.00 \times 10^8 \text{ m/s}} = 3.33 \times 10^{-7} \text{ T} = \boxed{0.333 \mu\text{T}}$$

$$(b) \quad \lambda = \frac{2\pi}{k} = \frac{2\pi}{1.00 \times 10^7 \text{ m}^{-1}} = \boxed{0.628 \mu\text{m}}$$

$$(c) \quad f = \frac{c}{\lambda} = \frac{3.00 \times 10^8 \text{ m/s}}{6.28 \times 10^{-7} \text{ m}} = \boxed{4.77 \times 10^{14} \text{ Hz}}$$

$$34.11 \quad S = I = \frac{U}{At} = \frac{Uc}{V} = uc$$

$$\frac{\text{Energy}}{\text{Unit Volume}} = u = \frac{I}{c} = \frac{1000 \text{ W/m}^2}{3.00 \times 10^8 \text{ m/s}} = \boxed{3.33 \mu\text{J/m}^3}$$

$$34.13 \quad r = (5.00 \text{ mi})(1609 \text{ m/mi}) = 8.04 \times 10^3 \text{ m}$$

$$S = \frac{\bar{P}}{4\pi r^2} = \frac{250 \times 10^3 \text{ W}}{4\pi(8.04 \times 10^3 \text{ m})^2} = \boxed{307 \mu\text{W/m}^2}$$

- 34.30 (a) If  $\mathcal{P}_S$  is the total power radiated by the Sun, and  $r_E$  and  $r_M$  are the radii of the orbits of the planets Earth and Mars, then the intensities of the solar radiation at these planets are:

$$I_E = \frac{\mathcal{P}_S}{4\pi r_E^2} \quad \text{and} \quad I_M = \frac{\mathcal{P}_S}{4\pi r_M^2}$$

$$\text{Thus,} \quad I_M = I_E \left( \frac{r_E}{r_M} \right)^2 = (1340 \text{ W/m}^2) \left( \frac{1.496 \times 10^{11} \text{ m}}{2.28 \times 10^{11} \text{ m}} \right)^2 = \boxed{577 \text{ W/m}^2}$$

- (b) Mars intercepts the power falling on its circular face:

$$\mathcal{P}_M = I_M (\pi R_M^2) = (577 \text{ W/m}^2) \pi (3.37 \times 10^6 \text{ m})^2 = \boxed{2.06 \times 10^{16} \text{ W}}$$

- (c) If Mars behaves as a perfect absorber, it feels pressure  $P = \frac{S_M}{c} = \frac{I_M}{c}$

$$\text{and force } F = PA = \frac{I_M}{c} (\pi R_M^2) = \frac{\mathcal{P}_M}{c} = \frac{2.06 \times 10^{16} \text{ W}}{3.00 \times 10^8 \text{ m/s}} = \boxed{6.87 \times 10^7 \text{ N}}$$

- (d) The attractive gravitational force exerted on Mars by the Sun is

$$F_g = \frac{GM_S M_M}{r_M^2} = \frac{(6.67 \times 10^{-11} \text{ N} \cdot \text{m}^2/\text{kg}^2) (1.991 \times 10^{30} \text{ kg}) (6.42 \times 10^{23} \text{ kg})}{(2.28 \times 10^{11} \text{ m})^2} = 1.64 \times 10^{21} \text{ N}$$

which is  $\sim 10^{13}$  times stronger than the repulsive force of (c).

34.31 (a) The total energy absorbed by the surface is

$$U = \left(\frac{1}{2}I\right)At = \left[\frac{1}{2}\left(750 \frac{\text{W}}{\text{m}^2}\right)\right](0.500 \times 1.00 \text{ m}^2)(60.0 \text{ s}) = \boxed{11.3 \text{ kJ}}$$

(b) The total energy incident on the surface in this time is  $2U = 22.5 \text{ kJ}$ , with  $U = 11.3 \text{ kJ}$  being absorbed and  $U = 11.3 \text{ kJ}$  being reflected. The total momentum transferred to the surface is

$$p = (\text{momentum from absorption}) + (\text{momentum from reflection})$$

$$p = \left(\frac{U}{c}\right) + \left(\frac{2U}{c}\right) = \frac{3U}{c} = \frac{3(11.3 \times 10^3 \text{ J})}{3.00 \times 10^8 \text{ m/s}} = \boxed{1.13 \times 10^{-4} \text{ kg} \cdot \text{m/s}}$$

34.43 (a)  $f\lambda = c$  gives  $(5.00 \times 10^{19} \text{ Hz})\lambda = 3.00 \times 10^8 \text{ m/s}$ :  $\lambda = 6.00 \times 10^{-12} \text{ m} = 6.00 \text{ pm}$

(b)  $f\lambda = c$  gives  $(4.00 \times 10^9 \text{ Hz})\lambda = 3.00 \times 10^8 \text{ m/s}$ :  $\lambda = 0.075 \text{ m} = 7.50 \text{ cm}$

\*34.58 The area over which we model the antenna as radiating is the lateral surface of a cylinder,

$$A = 2\pi r\ell = 2\pi(4.00 \times 10^{-2} \text{ m})(0.100 \text{ m}) = 2.51 \times 10^{-2} \text{ m}^2$$

(a) The intensity is then:  $S = \frac{\mathcal{P}}{A} = \frac{0.600 \text{ W}}{2.51 \times 10^{-2} \text{ m}^2} = \boxed{23.9 \text{ W/m}^2}$

(b) The standard is:  $0.570 \frac{\text{mW}}{\text{cm}^2} = 0.570 \left(\frac{\text{mW}}{\text{cm}^2}\right) \left(\frac{1.00 \times 10^{-3} \text{ W}}{1.00 \text{ mW}}\right) \left(\frac{1.00 \times 10^4 \text{ cm}^2}{1.00 \text{ m}^2}\right) = 5.70 \frac{\text{W}}{\text{m}^2}$

While it is on, the telephone is over the standard by  $\frac{23.9 \text{ W/m}^2}{5.70 \text{ W/m}^2} = \boxed{4.19 \text{ times}}$

34.64 Think of light going up and being absorbed by the bead which presents face area  $\pi r_b^2$ .

If we take the bead to be perfectly absorbing, the light pressure is  $P = \frac{S_{\text{av}}}{c} = \frac{I}{c} = \frac{F_{\ell}}{A}$

(a)  $F_{\ell} = F_g$  so  $I = \frac{F_{\ell}c}{A} = \frac{F_g c}{A} = \frac{m g c}{\pi r_b^2}$

From the definition of density,  $\rho = \frac{m}{V} = \frac{m}{\frac{4}{3}\pi r_b^3}$  so  $\frac{1}{r_b} = \left(\frac{4}{3}\pi\rho/m\right)^{1/3}$

Substituting for  $r_b$ ,  $I = \frac{m g c}{\pi} \left(\frac{4\pi\rho}{3m}\right)^{2/3} = g c \left(\frac{4\rho}{3}\right)^{2/3} \left(\frac{m}{\pi}\right)^{1/3} = \boxed{\frac{4\rho g c}{3} \left(\frac{3m}{4\pi\rho}\right)^{1/3}}$

(b)  $\mathcal{P} = IA = \boxed{\frac{\pi r_b^2 4\rho g c}{3} \left(\frac{3m}{4\pi\rho}\right)^{1/3}}$