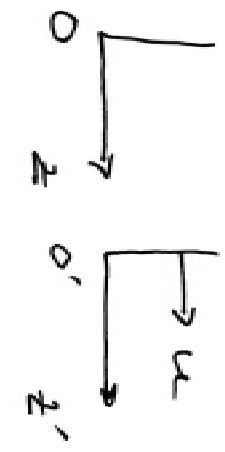


Relativity Review

To know:

- Inertial Reference Frames
 - $\Rightarrow$ Physics is the Same
- Lorentz transforms
 - Position and time (Contraction)
 - Velocity note: $v_1 \neq v_2$ not the same.
- Relativistic Doppler Shift
- Energy $\propto$ Momentum
- Photons $\propto$ Equivalence Principle

Lorentz transformations



$$x' = x ; \quad y' = y$$

$$z' = \gamma(z - ut) ; \quad t' = \gamma(t - \beta \frac{z}{c})$$

where $\gamma = \frac{1}{\sqrt{1 - \beta^2}}$; $\beta = \frac{u}{c}$

For the inverse transforms $x' \rightarrow x$

let $u \rightarrow -u$ and switch ' and un'

For the velocities

$$v_x' = \frac{v_x}{\gamma(1 - \beta \frac{v_x}{c})} ; \quad v_y' = \frac{v_y}{\gamma(1 - \beta \frac{v_x}{c})}$$

$$v_z' = \frac{v_z - u}{1 - \beta \frac{v_z}{c}}$$

Again to go from ' to un',
switch ' and un' and let
 $u \rightarrow -u$

Some Consequences

- Lorentz Contraction
- Time Dilation
- Doppler Shift

$$L_{\text{moving}} = \frac{L_{\text{rest}}}{\gamma}$$

$$\Delta t_{\text{moving}} = \gamma \Delta t_{\text{rest}}$$

$$\omega_{\text{moving}} = \omega_{\text{rest}} \sqrt{\frac{1 \pm \beta}{1 \mp \beta}}$$

where $+$ sign approaching
 $-$ sign receding

$$\lambda_{\text{moving}} = \lambda_{\text{rest}} \sqrt{\frac{1 \mp \beta}{1 \pm \beta}}$$

Consequence Continued

• Relativistic Energy

- Rest energy of massive particle

$$mc^2$$

- Total Energy of massive particle

$$\gamma mc^2$$

- Kinetic Energy of massive particle

$$\gamma mc^2 - mc^2 = mc^2(\gamma - 1)$$

- Relativistic Momentum

- massive particle

$$\gamma m v$$

- Energy & momentum for massless particle
like the photon

$$\begin{aligned} \xrightarrow{\text{total energy!}} E &= pc = h\nu = \frac{hc}{\lambda} \\ \Rightarrow p &= \frac{h}{\lambda} \end{aligned}$$

where h = Planck's Constant

For massive particle we have
another expression for the
total Energy

$$\underline{E = \sqrt{m^2 c^4 + p^2 c^2}}$$

must add the rest
energy to the energy
due to its motion

Quantum

To know:

- Bohr Atom - 1-electron atom
- | | |
|-------------------------|-------------------------|
| • de Broglie wavelength | } wave nature of matter |
| • Uncertainty Principle | |
- Schrodinger { equation (time-independent)
- Wave function, probability density
- Example of infinite square well potential

The Photon

Quantum Mechanics (Photoelectric Effect, Compton

Scattering, etc.) tell us light comes in packages
of energy of quantity $h\nu$. How do we

reconcile this with our notion of
intensity in a beam of light?

$$I = \frac{\text{Power}}{\text{unit area}} = \frac{\text{Energy/unit time}}{\text{unit area}}$$

$\Rightarrow$ introduce the notion of

Photon Flux, η ,

Photon per unit time per unit area

$$I \sim \frac{n}{\Delta t \cdot A} h\nu \quad \begin{array}{l} \Delta t = \text{length of time} \\ A = \text{area} \end{array}$$

total Energy in the beam

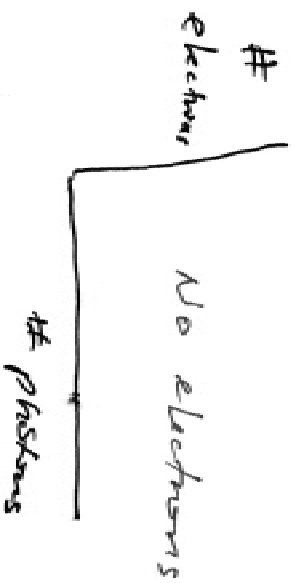
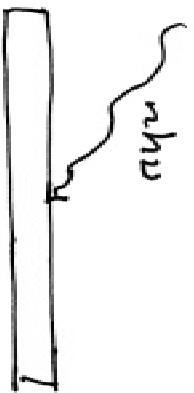
$$E = I \cdot \Delta t \cdot A = n h\nu \quad \begin{array}{l} \text{Amount of energy} \\ \text{in each photon} \end{array}$$

↓
Package

of photons

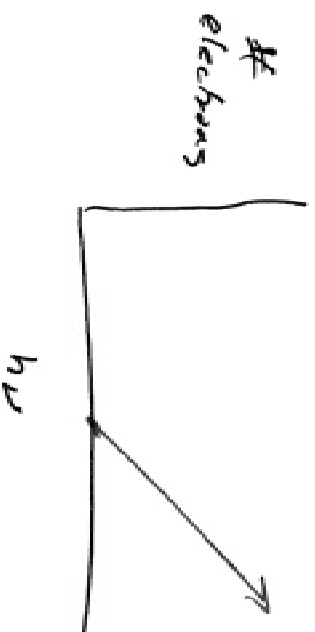
⇒ there are two aspects to the energy of light — the amount of energy in each unit & the number of units.

This means



If we increase the energy by raising ν , we will eventually reach a threshold where electrons are liberated.

If we increase the energy by raising the energy (and thus the intensity) by increasing the # of photons no electrons are liberated if $h\nu < \text{work function}$



Quantum Mechanics

$$h \quad \frac{1}{2} \quad h = \frac{h}{2\pi}$$

Should know in units of J-s & eV-s

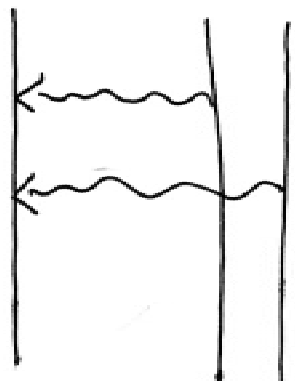
has units of energy-time

1- electron atom Rydberg formula tells us the wavelength of the light due to transitions between bound (stationary) energy levels

$n = \infty$  electron free

$n = 3$
 $\vdots$
 $n = 2$

$n = 1$



$$\frac{1}{\lambda_{nm}} = R_{\infty} \left(\frac{1}{n^2} - \frac{1}{m^2} \right)$$

Note $\frac{1}{\lambda_{n\infty}} = \frac{R_{\infty}}{n^2}$

so for $n=1$ $\lambda_{1\infty} = \frac{1}{R_{\infty}}$

In the Bohr model

$$E_n = -\frac{1}{2} \left(\frac{e^2}{4\pi\epsilon_0} \right) \frac{1}{n^2 a_0}$$
$$= -hc \frac{R_\infty}{n^2}$$

for the levels of
hydrogen

$a_0 \approx$ Bohr radius

$$\underbrace{\Delta E = E_{n'} - E_n = \frac{hc}{\lambda_{n'}} = hc R_\infty \left(\frac{1}{n^2} - \frac{1}{n'^2} \right)}$$

The amount of energy required to excite

an electron for level n to a higher level n'

Note, the larger ΔE is the smaller $\lambda_{n'}$ is!

Wave nature of matter.

As with light de Broglie

$$\lambda = \frac{h}{p} = \frac{h}{mv} \quad \text{non relativistically}$$

W/ this all interference & diffraction
you learned for light now applies
to matter.

As a result of the wave nature we have
the uncertainty principle

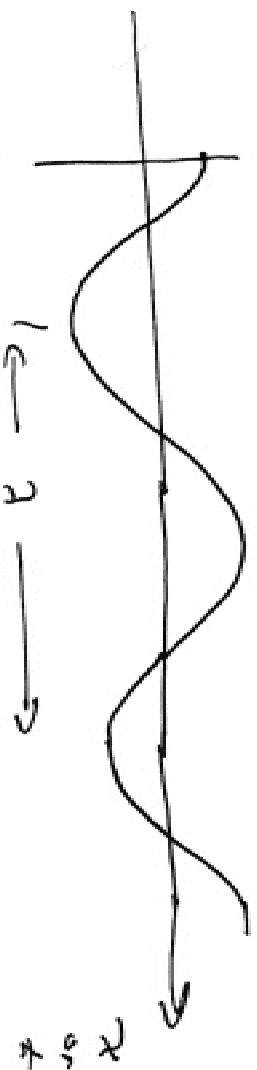
$$\Delta x \Delta p \geq \frac{\hbar}{2} \quad \text{and} \quad \Delta E \Delta t \geq \frac{\hbar}{2}$$

What does this mean?

You cannot determine both the position and the momentum to an arbitrary accuracy. Consider a wave of a single frequency.

$$\psi(x, t) = A \cos(kx - \omega t)$$

This describes a plane wave. We know very well $h \pm \omega$ but as a consequence the particle is everywhere at the same time



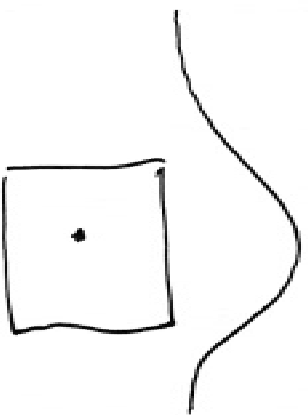
λ is known exactly $\Rightarrow$ k is known exactly

$$p = \hbar k \Rightarrow p \text{ known exactly}$$

but x runs from $-\infty$ to $+\infty$

we have no knowledge of where the particle is!

Now say


$$= \sum_{k_n} \cos(k_n x)$$

different
frequencies!

we know where the particle is $\Delta x = L$

in order to restrict the wave function to this box we had to add waves of different frequencies

$$\text{So } \Delta p \approx \frac{\hbar}{2\Delta x} \approx \frac{\hbar}{2L}$$

Time-independent Schrodinger Equation

$$\frac{d^2}{dx^2} \psi(x) = \frac{2m}{\hbar^2} [V(x) - E] \psi(x)$$

→ Solve for $\psi(x)$ for given $V(x)$

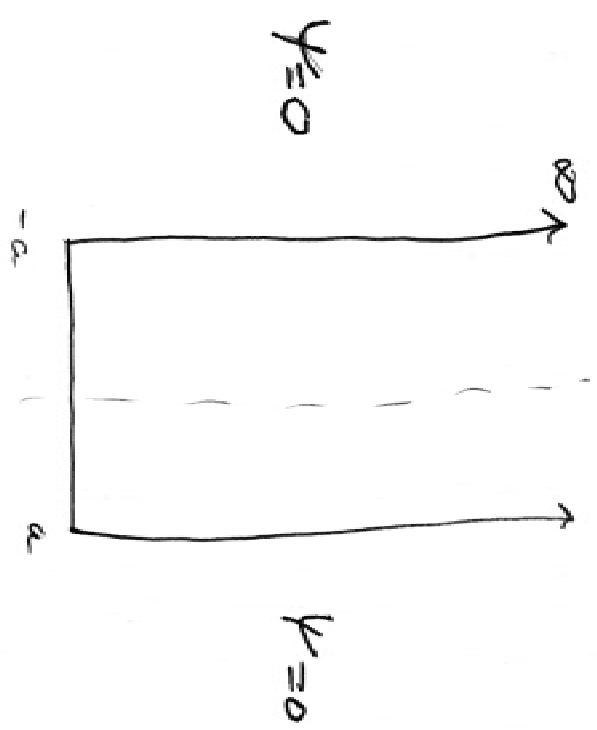
When $V(x) = 0$

$$\frac{d^2}{dx^2} \psi(x) = -\frac{2mE}{\hbar^2} \psi(x)$$

Solution $\psi(x) = A \cos kx + B \sin kx$

where $k = \frac{\sqrt{2mE}}{\hbar}$

Example Particle in infinite square well potential



inside $\psi(x) = A \cos kx + B \sin kx$

$\psi(a) = \psi(-a) = 0$

Even solutions

$A \cos ka = 0 \iff \cos$ an even function,

$ka = \frac{n\pi}{2} \quad n = 1, 3, 5, \dots$

$k = \frac{n\pi}{2a}$

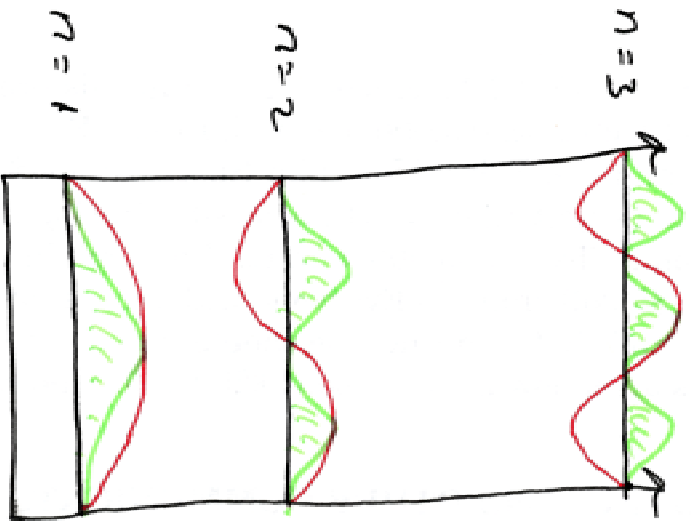
Odd solutions

$B \sin ka = 0$

$k = \frac{n\pi}{2a} \quad n = 2, 4, 6, \dots$

Kinetic Energy is now Quantized

$$E_n = \frac{\hbar^2 k_n^2}{2m} = n^2 \frac{\pi^2 \hbar^2}{8ma^2} = \frac{n^2 h^2}{32ma^2}$$



$\psi(x)$ the amplitude

$|\psi(x)|^2$ the probability density

$$\begin{aligned} \text{Prob} &= \int_{-a}^0 |\psi(x)|^2 dx \\ &= 50\% \end{aligned}$$

Probability for finding the particle between $-a$ and 0 .