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February 13, 2013 Physics 132 Prof. E. F. Redish

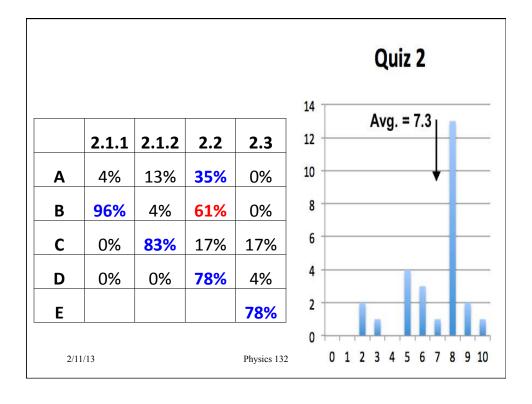
Theme Music: Doris Day

Que Sera, Sera

Cartoon: Bill Watterson

Calvin & Hobbes

HE SAYS HE HASNIT USED A SLIDE RULE SINCE, BECAUSE IN SCHOOL, THEY TAUSH HIM TO DO MATH ON A SLIDE RULE. CALCULATOR THAT CAN DO MORE FUNCTIONS THAN HE COUT FIRST OUT FIRST O



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Foothold ideas: Exponents and logarithms



- Power law: $f(x) = x^2$ $g(x) = Ax^7$ a variable raised to a fixed power.
- Exponential: $f(x) = e^x$ $g(N) = 2^N$ $h(z) = 10^z$ a fixed constant raised to a variable power.
- Logarithm: the inverse of the exponential.

$$x = e^{\ln(x)} \qquad x = \ln(e^x)$$
$$y = 10^{\log(y)} \qquad y = \log(10^y)$$

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$$\log(2) = 0.3010$$

$$\log(e) = 0.4343$$

$$2^{N} = (10^{0.3010})^{N} \approx 10^{0.3N}$$

$$e^{x} = (10^{0.4343})^{x} \approx 10^{0.4x}$$

$$2^{N} = B$$

$N\log 2 = \log B \Rightarrow N = \frac{\log B}{\log 2}$

Foothold ideas: Entropy



- Entropy an extensive measure of how well energy is spread in a system.
- **■** Entropy measures
 - The number of microstates in a given macrostate $S = k_B \ln(W)$
 - The amount that the energy of a system is spread among the various degrees of freedom
- Change in entropy upon heat flow

$$\Delta S = \frac{Q}{T}$$

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Foothold ideas: Transforming energy



■ Internal energy: thermal plus chemical

 ΔU

- Enthalpy: $\Delta H = \Delta U + p\Delta V$ internal plus amount needed to make space at constant p
- Gibbs free energy: $\Delta G = \Delta H T \Delta S$ enthalpy minus amount associated with raising entropy of the rest of the universe due to energy dumped
- A process will go spontaneously if $\Delta G < 0$.

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Spontaneity...

$$\Delta G = \Delta H - T\Delta S$$

$$-T\Delta S_{\text{total}} - T\Delta S_{\text{surroundings}} T\Delta S_{\text{system}}$$

The sign of the Gibbs Free Energy change indicates spontaneity!

$$\Delta G < 0 \rightarrow \Delta S_{\text{total}} > 0 \rightarrow \text{spontaneous}$$

 $\Delta G > 0 \rightarrow \Delta S_{\text{total}} < 0 \rightarrow \text{not spontaneous}$

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Foothold ideas: Energy distribution

- Due to the randomness of thermal collisions, ever in (local) thermal equilibrium a range of energy is found in each degree of freedom.
- The probability of finding an energy E is proportional to the Boltzmann factor

$$P(E) \propto e^{-E/k_B T}$$
 (for one DoF)
 $P(E) \propto e^{-E/RT}$ (for one mole)

■ At 300 K, $k_{\rm B}T \sim 1/40 \text{ eV}$ $N_{\rm A}k_{\rm B}T = RT \sim 2.4 \text{ kJ/mol}$ Physics 132

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