

Homework Solutions, Physics 117

Home Work Problem Set # 3

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Solutions by ✓

Ch 4: Q 9, 23; Ex 7, 11 // Ch 5: Q: 3, 9, 13; Ex 3, 11, 19.

4: Q 9

- (a) Speeding up & moving in a straight line, because \vec{a} is \parallel to \vec{v}
 (b) Speeding up & turning left, because component of $\vec{a} \perp \vec{v}$ is directed upward, and generates \vec{v} component to the left of forward.

4: Q 23

The same time, since the vertical and horizontal motions are separate and independent, although simultaneous.
 [But if the rock were thrown so fast horizontally that the curvature of the moon's surface became important, see pages 79-81]

4: Ex 7.

$$a_c = \frac{v^2}{r} = \frac{(20)^2}{50} = \boxed{8 \text{ m/sec}^2}; F = ma = (120)(8) = \boxed{960 \text{ N.}}$$

4: Ex 11.

Neglecting Air resistance, $v_{\text{HORIZ}} = \boxed{22 \text{ m/sec}}$ at $t=1 \text{ sec}$.
 $\Delta v_{\text{VERT}} = v_0 - gt = 14 - 10 \cdot 1 = \boxed{4 \text{ m/sec}}$ at $t=1 \text{ sec}$.
 because the acceleration is vertically downward.

5: Q 3:

The force that Earth exerts on moon is SAME (in magnitude)
 as force " Moon " " EARTH, by (NIII).

5: Q 9:

$$F_G = \frac{GMm}{r^2} \text{ on Earth } F_G = 800 \text{ N} = \frac{GM_E m_{\text{SP}}}{R_E^2}$$

$$\text{on planet MUSI3, } F_G = \frac{GM_E \cdot m_{\text{SP}}}{R_{\text{MUSI3}}^2} = \frac{800}{4} = \boxed{200 \text{ N} = F_G^{\text{MUSI3}}}$$

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5: Q 13 $a = \frac{F_g}{m_{app}} = \frac{G M_E m_{app}}{(R_E)^2} = g = \frac{10 \text{ m}}{\text{sec}^2} \text{ ON EARTH}$

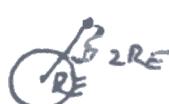
in our universe, where $G = 6.67 \times 10^{-11} \frac{\text{N m}^2}{\text{kg}^2}$.

In the parallel universe with M_E and m_{app} and R_E the same, the acceleration $a_p = g_p = 20 \text{ m/sec}^2 = \frac{G_p \cdot M_E m_{app}}{m_{app} \cdot (R_E)^2}$

Then $\frac{g_p}{g} = \frac{\frac{G_p M_E m_{app}}{m_{app} (R_E)^2}}{\frac{G M_E m_{app}}{m_{app} (R_E)^2}} = \frac{G_p}{G} = \frac{20}{10} = 2$

Thus, in the parallel universe, the Gravitational Constant has a value twice that of G in our universe: $G_p = \boxed{2 \times G} = 13.34 \times 10^{-11} \frac{\text{Nm}^2}{(\text{kg})^2}$

5 Ex 3 : $\frac{F_g}{m} = \text{accel due to gravity} = \frac{G M_E m}{r^2 (3R_E)^2} = \frac{1}{9} \frac{G M_E}{(R_E)^2} = \frac{g}{9}$

 Since distance = $3R_E$ to center's center.

$$= \boxed{\frac{1.1 \text{ m}}{\text{sec}^2}}$$

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By <u>Q1</u>
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5: Ex 11 $F = \frac{GM_E m}{r^2} = 800 N$ when $m = 320 \text{ kg}$.
 where $\frac{GM_E}{(R_E)^2} = g = 10 \text{ m/sec}^2 = 10 \frac{\text{N}}{\text{kg}}$

$$F = 800 N = \frac{GM_E m}{(R_E)^2 (T/RE)^2} = \frac{(g) 320 N}{(T/RE)^2} \quad \& \quad R_E = 6.4 \times 10^6 \text{ m.}$$

$$(T/RE)^2 = \frac{10 \cdot 320}{800} = 4.0 \Rightarrow T = \sqrt{4} T_{RE} = \text{distance} = \frac{2 R_E}{12.8 \times 10^6 \text{ m}}$$

$$\text{and ALTITUDE} = T - R_E = 1.0 R_E = 6.4 \times 10^6 \text{ m.}$$

[Using $R_E = 6.4 \times 10^6 \text{ m}$. Find the Radius of EARTH from inside back cover.]

5: Ex 19 If Orbit has $r = 6.6 R_E$ & $T = 24 \text{ hours}$, then

$$\text{satellite travels } 2\pi r = 2\pi (6.6) \times 6.4 \times 10^6 \text{ m} = [2.65 \times 10^9 \text{ m/day}]$$

$$\& \text{ its orbital velocity is } = \frac{2.65 \times 10^9 \text{ m}}{\text{day}} \times \frac{1 \text{ day}}{24 \text{ hrs}} \times \frac{1 \text{ hr}}{60 \text{ min}} \times \frac{1 \text{ min}}{60 \text{ sec}}$$

$$= 3.07 \times 10^3 \frac{\text{m}}{\text{sec}}$$

end of HW #3 solutions